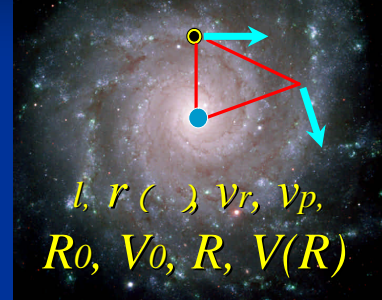


回転曲線および 近太陽円銀河定数 観測の最適化

祖父江義明
松井、中川、安藤、今井、中西、本間*, et al.
鹿児島大学, *NAO

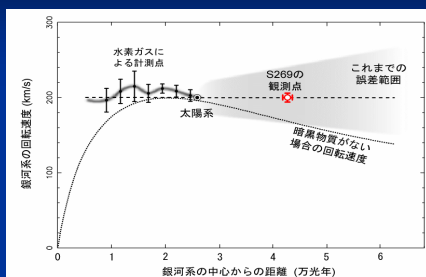
2008.10.9-10, VERA UM, Mitaka

・ 回転曲線

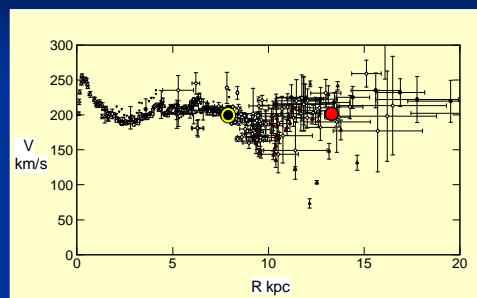


$l, r(\) v_r, v_p,$
 $R_0, V_0, R, V(R)$

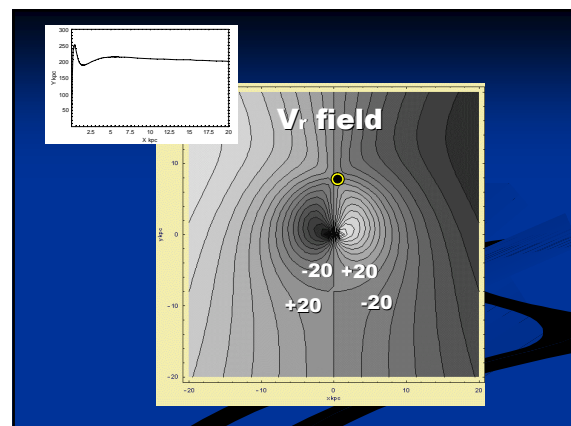
銀河系回転曲線 (Honma et al. 2007)



$V_0=200 \text{ km/s}, R_0=8 \text{ kpc}$
(Sofue, Honma, Omodaka, et al. 2008)



回転曲線 - 1 円軌道を仮定して [v_r, η] 測定から (従来の方法)



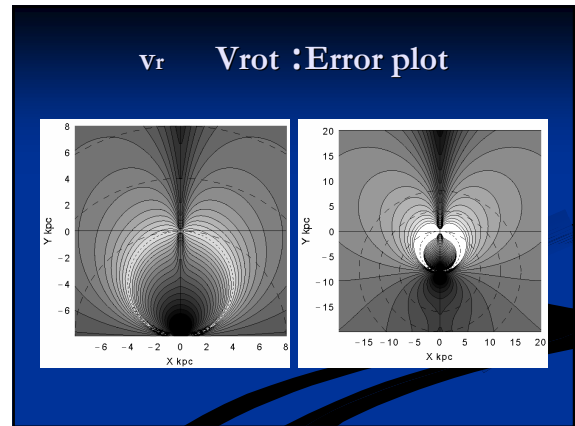
$$V_{rot}^{rad} = \frac{R}{R_0} \left(\frac{v_r}{\sin l} + V_0 \right)$$

$$\delta V_{rot}^{rad} = \sqrt{\delta V_{vr}^2 + \delta V_r^2}$$

$$\delta V_{vr} = \frac{\partial V}{\partial v_r} \delta v_r, \quad \delta V_r = \frac{\partial V}{\partial r} \delta r$$

$$\delta V_{rot}^{rad} = \sqrt{\left(\frac{R}{R_0 \sin l} \right)^2 \delta v^2 + \left(\frac{s V}{R^2} \right)^2 \delta r^2}$$

$V = V_0 = 200 \text{ km s}^{-1}$
 $\delta v = 2 \text{ km s}^{-1}$
 $\delta r/r = 0.1$



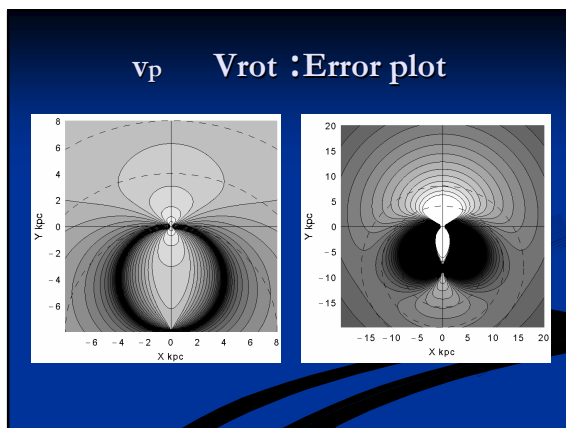
回転曲線 -2

円軌道を仮定して
 $[v_p, \eta]$ 測定から
 $(v_r$ が不定の場合)

$$V_{rot}^{prop} = -\frac{R}{s} (v_p + V_0 \cos l)$$

$$\delta V_{rot}^{prop} = \sqrt{\left(\frac{R}{s} \right)^2 \left[\delta v_p^2 + \left(\frac{r R_0^2 v_p \sin^2 l}{s R^2} \right)^2 \left(\frac{\delta r}{r} \right)^2 \right]}$$

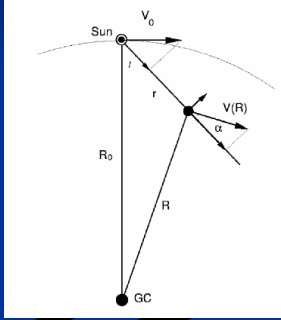
$V = V_0 = 200 \text{ km s}^{-1}$
 $\delta \mu = 0.21 \text{ mas y}^{-1}$
 $\delta v_p/r = 1 \text{ km s}^{-1} \text{ kpc}^{-1}$
 $\delta r/r = 0.1$



回転曲線 -3

究極のVERA法
円軌道を仮定せず
 v_p, v_r, η 測定から
直接V-vector

[v_p, v_r, \hat{r}] V vector



Velocity Vector from v_p, v_r , and \hat{r}

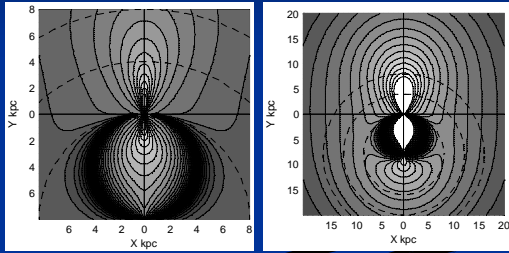
$$V = [(v_p + V_0 \cos l)^2 + (v_r - V_0 \sin l)^2]^{1/2}$$

$$\delta V = [\delta v_p^2 + \delta v_r^2]^{1/2} = \left[\left(\frac{\partial V}{\partial v_p} \delta v_p \right)^2 + \left(\frac{\partial V}{\partial v_r} \delta v_r \right)^2 \right]^{1/2}$$

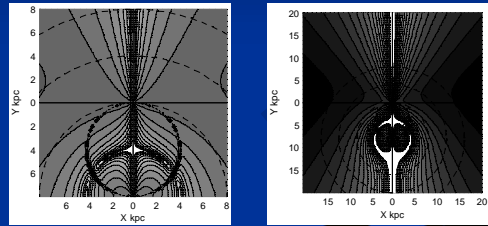
$$v_r = \left(V \frac{R_0}{R} - V_0 \right) \sin l.$$

$$v_p = \frac{r}{R} (R_0 \cos l - r) - V_0 \cos l,$$

v_r, v_p, \hat{r} V-vector: Error plot



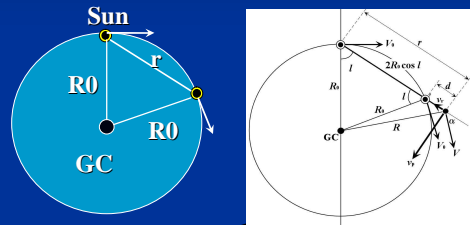
V-vector Error plot, $dv_r \gg dv_p$



銀河定数
近太陽円法

cf VERA KP Honma et al;
cf 安藤、中川、et al

R_0, V_0 決定
真太陽円 近太陽円



真太陽円

$$r = 2R_0 \cos l,$$

$$v_p = -2V_0 \cos l.$$

$$R_0 = \frac{r}{2 \cos l}$$

$$V_0 = -\frac{v_p}{2 \cos l}$$

$$R_0 = \frac{r}{2 \cos l} \left(1 + \frac{v_r}{v_p \sin l} \right) = \frac{r}{2 \cos l} \left(1 - \frac{d}{r} \right)$$

$$V_0 = -\frac{v_p}{2 \cos l} \left(1 - \frac{d}{r} - 2 \frac{v_r \cos^2 l}{v_p \sin l} \right) = -\frac{v_p}{2 \cos l} \left(1 - \frac{d}{r} \right) + v_r \cot l$$

$$\delta R_0 = \frac{1}{2 \cos l} \left[\delta r^2 + \left(\frac{\delta v_r}{v_p \sin l} \right)^2 r^2 \right]^{1/2}$$

$$\delta V_0 = \frac{1}{2 \cos l} \left[\left(\frac{1}{\sin l} - \frac{2 \cos^2 l}{v_p \sin l} \right)^2 v_p^2 \delta v_r^2 + \delta v_p^2 \right]^{1/2}$$

近太陽円

VERA, ON2, $l=76$ deg (Ando, Nakagawa, et al. 2008)

Table 1. VERA Determination of the Galactic Constants R_0 and V_0 by Solar Circle Method - Observations

	r kpc	δr kpc	v_r km s ⁻¹	δv_r km s ⁻¹	v_p km s ⁻¹	δv_p km s ⁻¹
VERA (wtd mean v_r)	9.34	±0.34	403	±5.15	-133.2	±6.66
VERA (mean v_r)	8.66	±4.66	250.76	±10.58	-133.2	±6.66
† G75.8+0.40 H92 α	8.29	±4.43	239.69	±13.80	-133.2	±6.66
† G75.8+0.35 H92 α	7.06	±4.43	202.98	±14.22	-133.2	±6.66

Preliminary.
参考!

VERA Determination of the Galactic Constants R_0 and V_0 by Solar Circle Method

	R_0 kpc	δR_0 kpc	V_0 km s ⁻¹	δV_0 km s ⁻¹
On-circle Values	9.34	-	271.27	-
VERA (wtd mean v_r)	8.66	±4.66	250.76	±46.0
VERA (mean v_r)	7.42	±5.36	213.71	±91.3
G75.8+0.40 H92 α	8.29	±4.43	239.69	±13.80
G75.8+0.35 H92 α	7.06	±4.43	202.98	±14.22

真太陽円仮定

近太陽円法

R_0, V_0 決定近太陽円法 最適化

