

ALMA workshopのお知らせ

ALMAとVLBIの連携に向けて

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AIM of this workshop

- ALMA sciences with the highest resolution
 - Comparable to VLBI (<10mas)
 - Expansion from VLBI to ALMA
 - Complement to VLBI
- Initial studies in the cycle 0
- Future work with the highest resolution
 - Full operation; longer baselines, higher frequency
- Start collaboration
 - Not only within radio/VLBI community
 - WITH wider field of astronomy

From the VLBI view

- Advantage of VLBI
 - Higher resolution in ALMA (<1 mas) 30 years ago!
 - Imaging of maser sources, AGN, and so on . . .
 - Proper motion measurements, astrometry
 - 3-dimensional velocity structure (in case of masers)
- Disadvantage of VLBI
 - Lowest (brightness) sensitivity ($>10^9\text{K}$)
 - Impossible to detect thermal dust continuum and molecular line emission
 - Hard to estimate physical parameters
- Complement by ALMA in the similar resolution

First meeting

- 2011 June 1 at NAOJ
- In particular for Cycle 0
- Program
 - Introduction; Hirota, Okumura, Nagai
 - Stellar Evolution; Kameno, Matsumoto, Imai
 - Galaxies and AGN; Hayashi, Doi, Nagai, Hada
 - ISM/SFRs; Motogi, Nagayama, Sugiyama, Sunada

<http://veraserver.mtk.nao.ac.jp/restricted/>

ALMAWS2011/index.html

Future

- **Workshop for each field will be scheduled**
 - Review the cycle 0 proposal
 - Discuss about cycle 1 proposal
 - Study more toward full operation
- **Please let me know your comments**
 - Preferred date, venue, scientific theme, possible invited reviewer/speaker, and more ...

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