

# A KVN-VERA Joint Imaging Observation of 44 GHz Methanol Maser toward G18.34+1.78

N. Matsumoto, T. Hirota, Mikyoung Kim, M. Honma (NAOJ),  
K. Sugiyama (Yamaguchi Uni.),  
Do-Young Byun, Taehyun Jung, Jongsoo Kim, Kee-Tae Kim (KASI),  
and VERA+KVN SFR sub-science WG

# Abstract

*The world's first*

## **VLBI imaging of 44GHz methanol maser!**

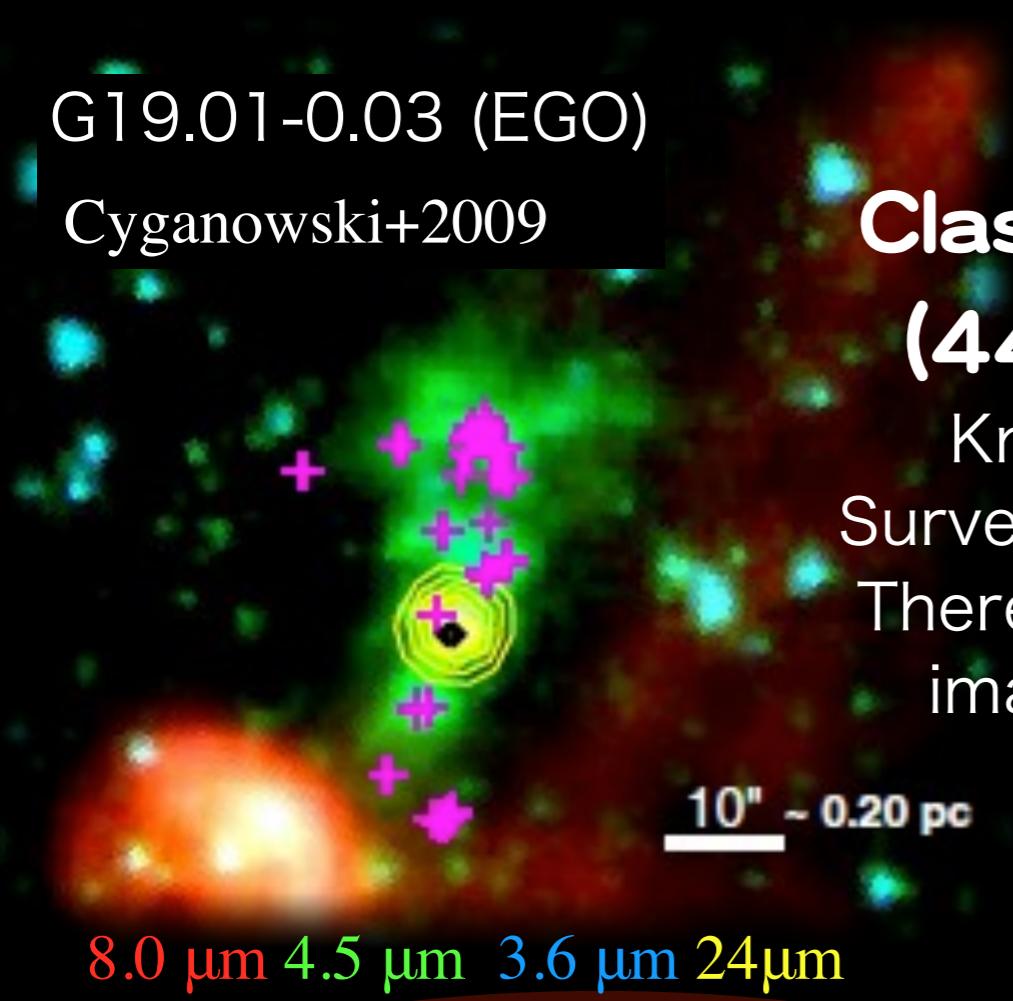
- We conducted first time observations of 44GHz methanol maser with KVN+VERA.
- Because of the extended distribution around 100 AU, it has been thought that detecting 44GHz methanol maser with VLBI is difficult so far.
- But, KVN+VERA array was able to detect 44GHz methanol maser toward G18.34+1.78 with short baselines and good UV-coverage/efficiency.  
→We were able to confirm the spot structure on images.

Intr.

# Methanol maser around Massive star forming regions

G19.01-0.03 (EGO)

Cyganowski+2009



VLBI can detect proper motions  
and spacial distributions in  
circumstellar gases with few  
milli arc second scale.

**Class I methanol maser  
(44GHz  $7_0-6_1A^+$  etc.)**

Known sources are limited.  
Surveys are currently in progress.  
There are dozens interferometer  
images, but no VLBI images.

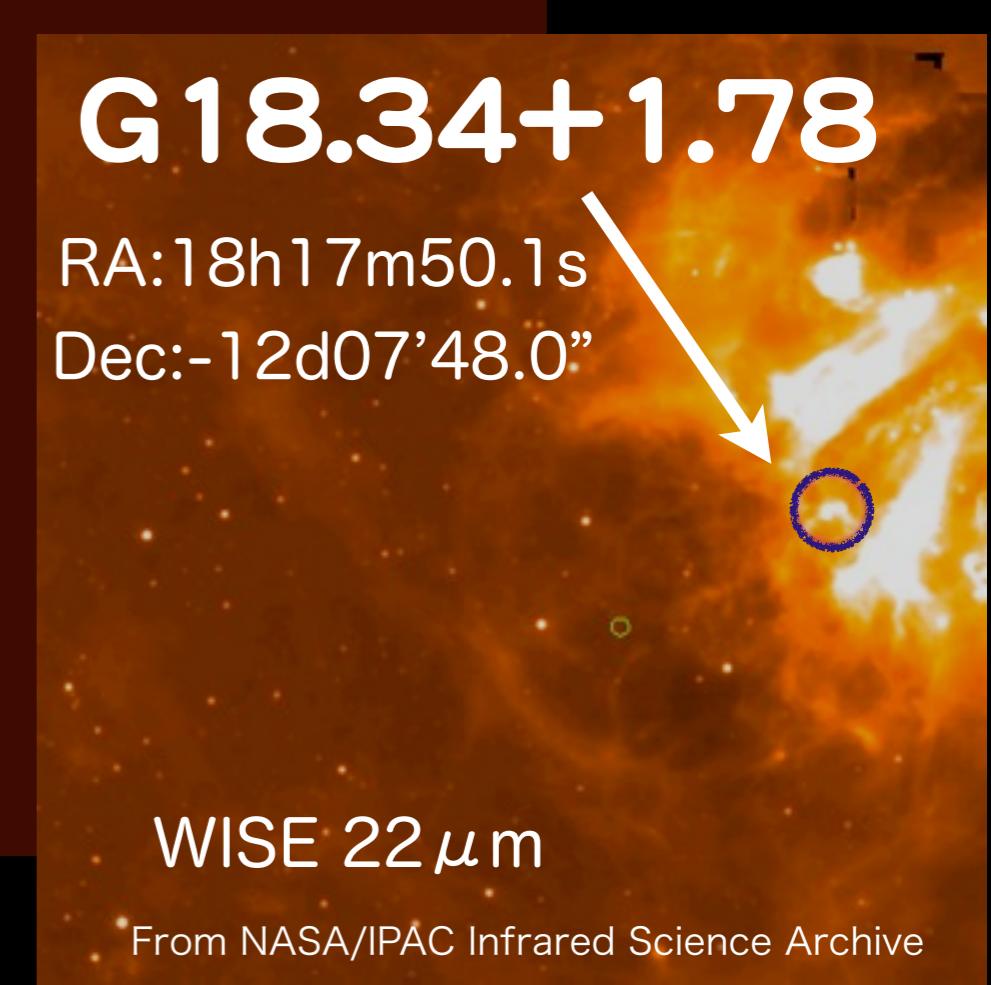
**Class II methanol maser  
(6.7GHz  $5_1-6_0A^+$  etc.)**

Over 1,000 sources have already  
detected by large survey projects.  
There are some statistical study  
with VLBI imaging.

**Obs.**

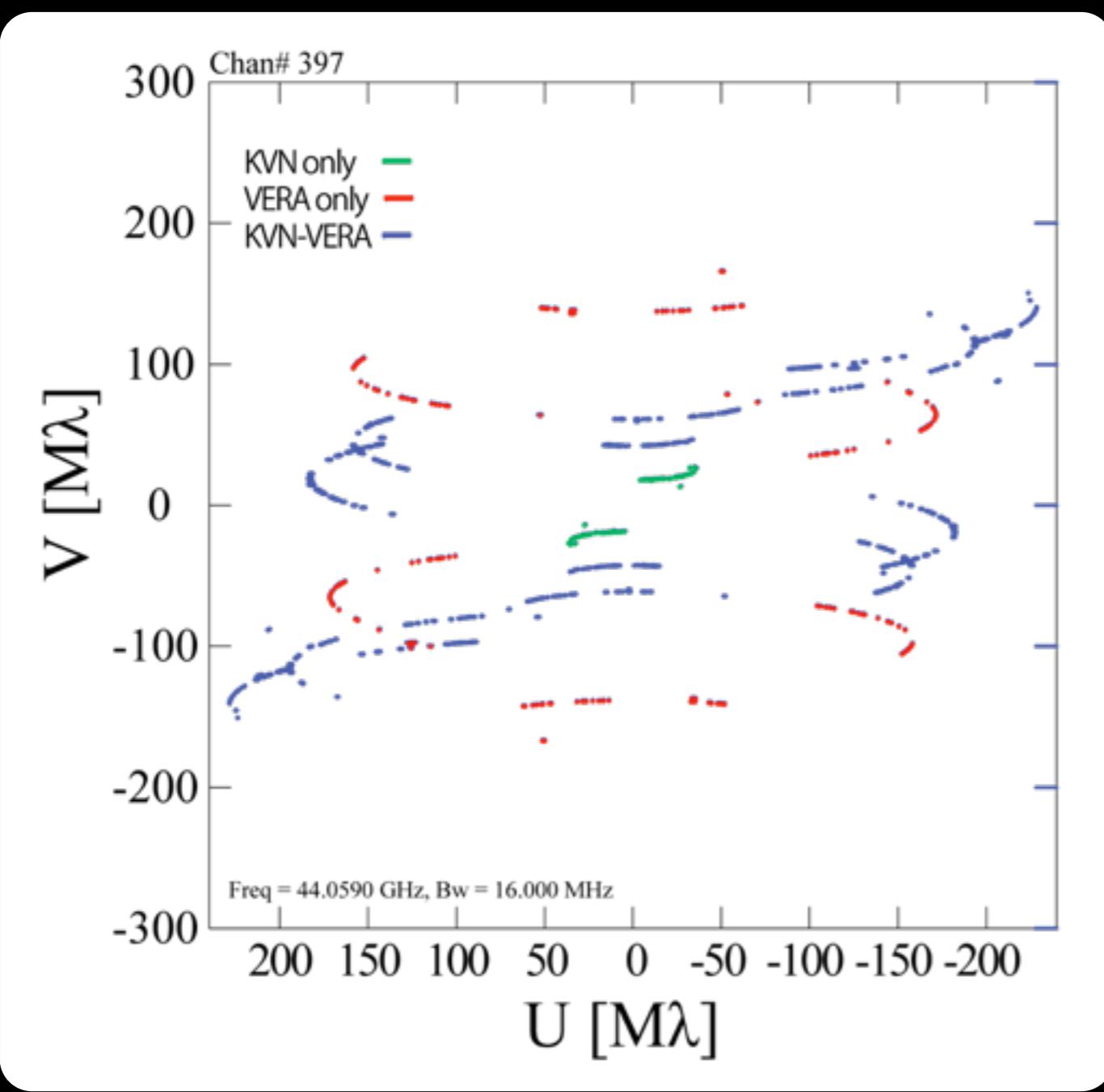
# 1st imaging observations of 44GHz methanol maser sources with KVN+VERA

- Date : Apr. 8, 2012, UT 17h - 24h
- Line : 44GHz ClassI methanol maser
- Target : G18.34+1.78
- $D_{\text{kin}}$  :  $\sim 2.7 \text{kpc}$
- Array: KVN(2)+VERA(3)
- Recording : DIR1000 128Mbps



# Results

## UV coverage is improved with KVN+VERA

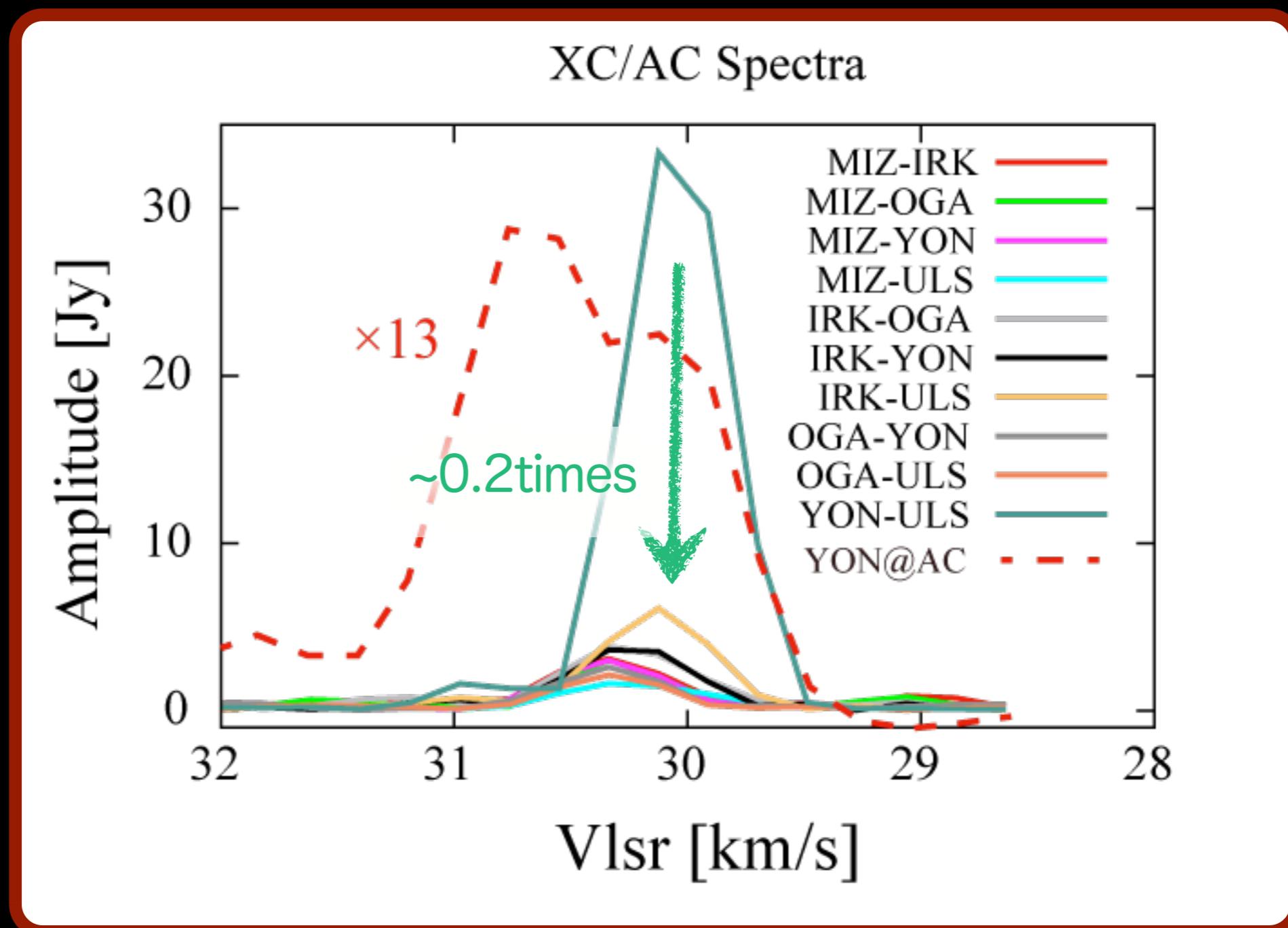


KVN+VERA  
observations produce a  
complementary effect.

We get a new window!!

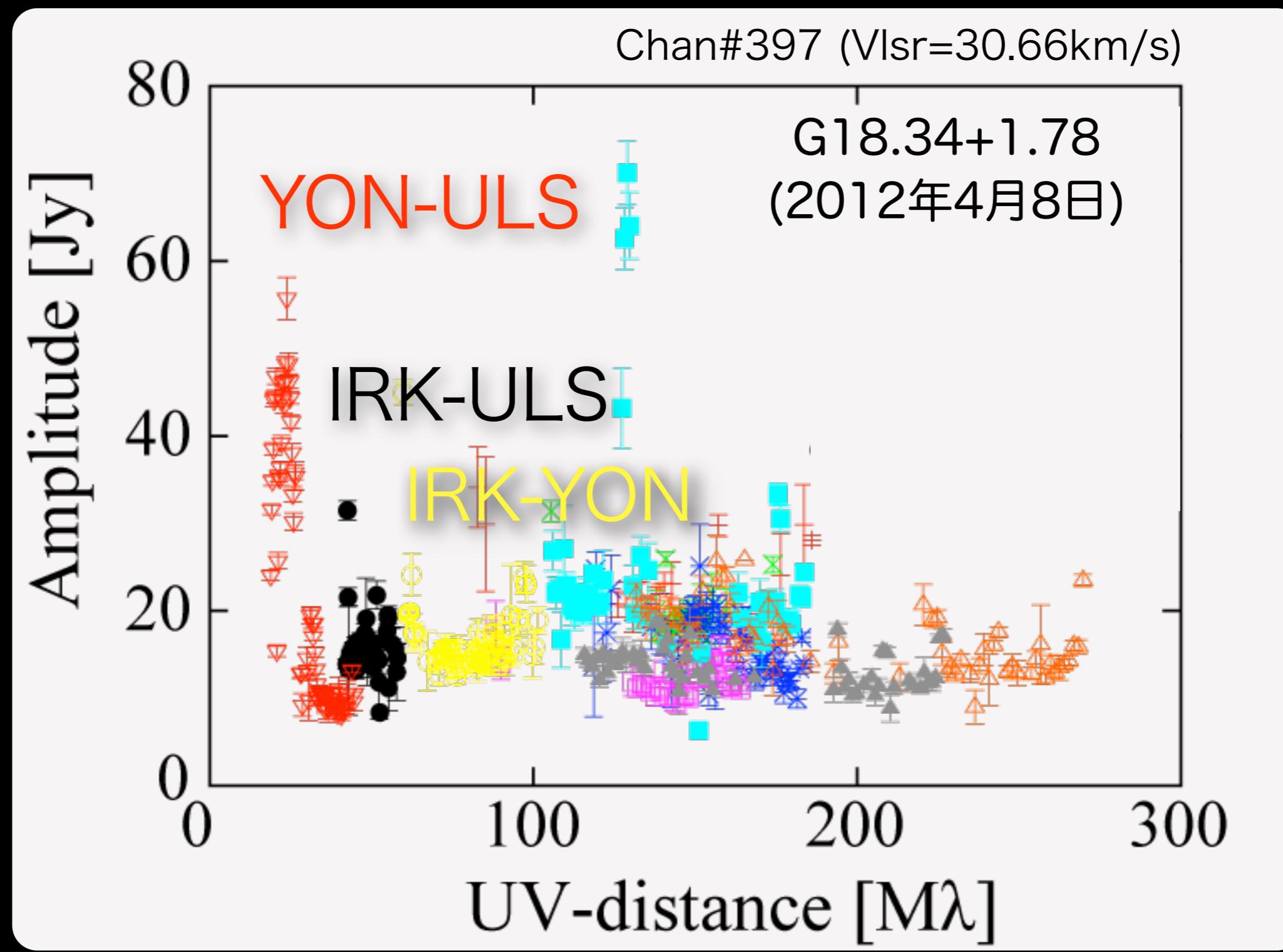
Apr. 8, 2012  
G18.34+1.78

# Highly resolved out



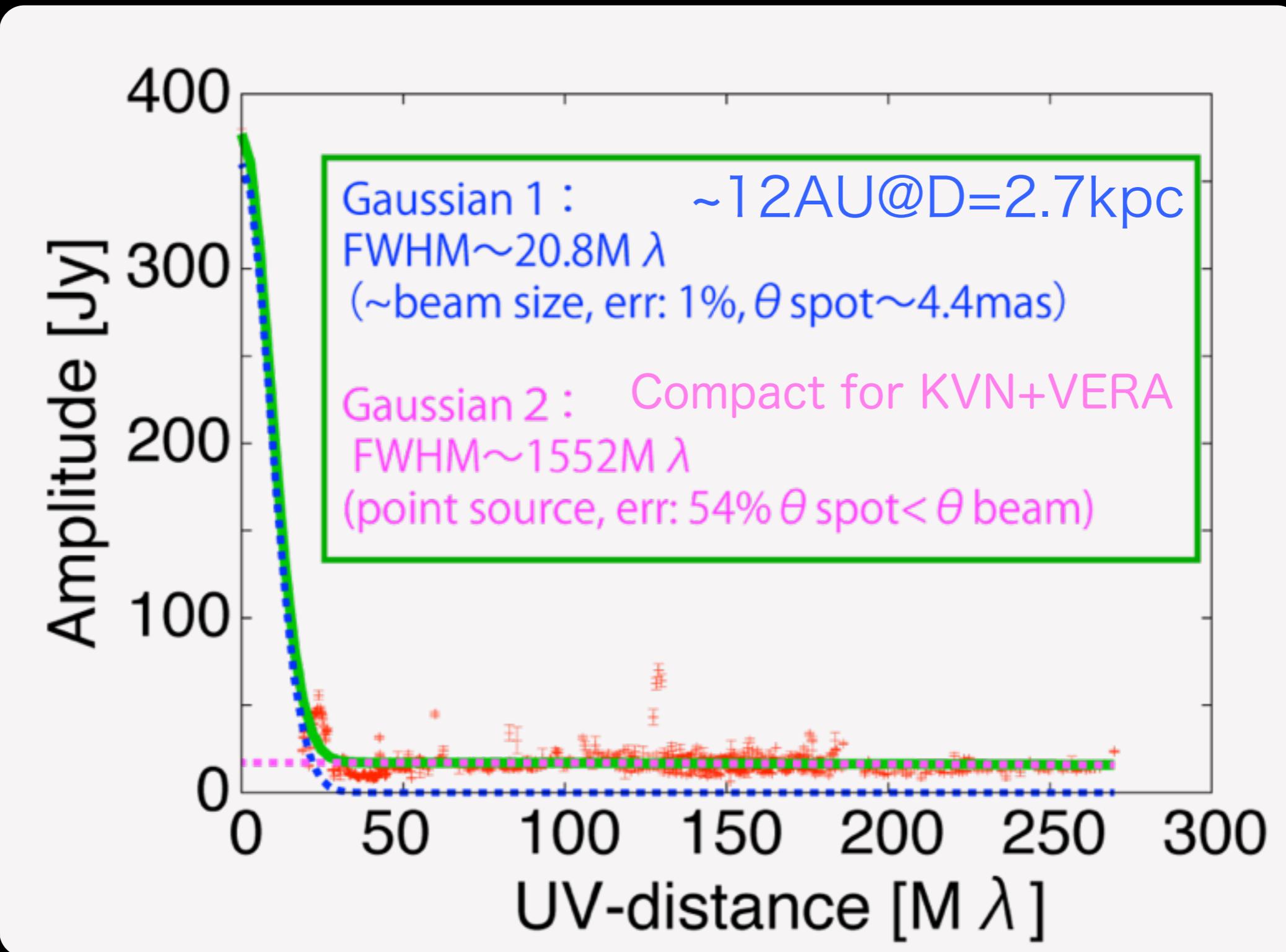
44GHz methanol maser toward G18.34+1.78  
(Apr. 8, 2012)

# Highly resolved out



# Results

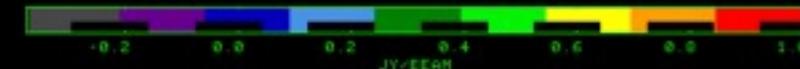
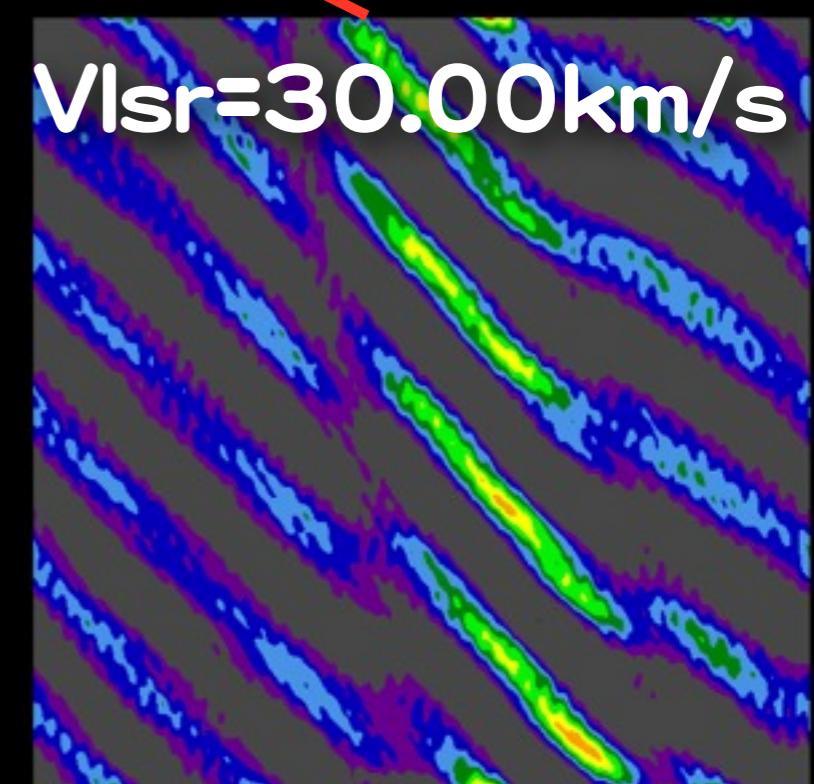
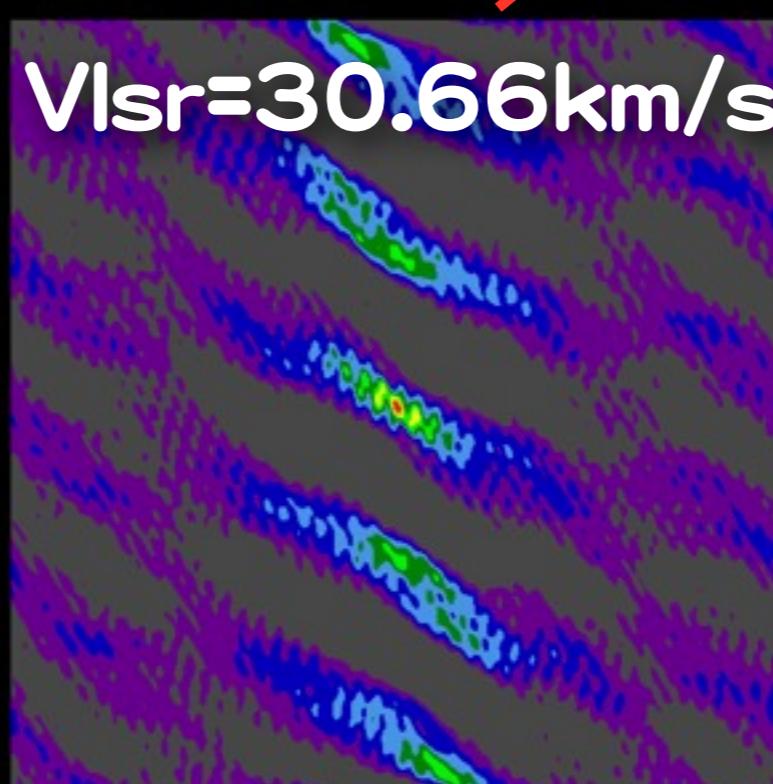
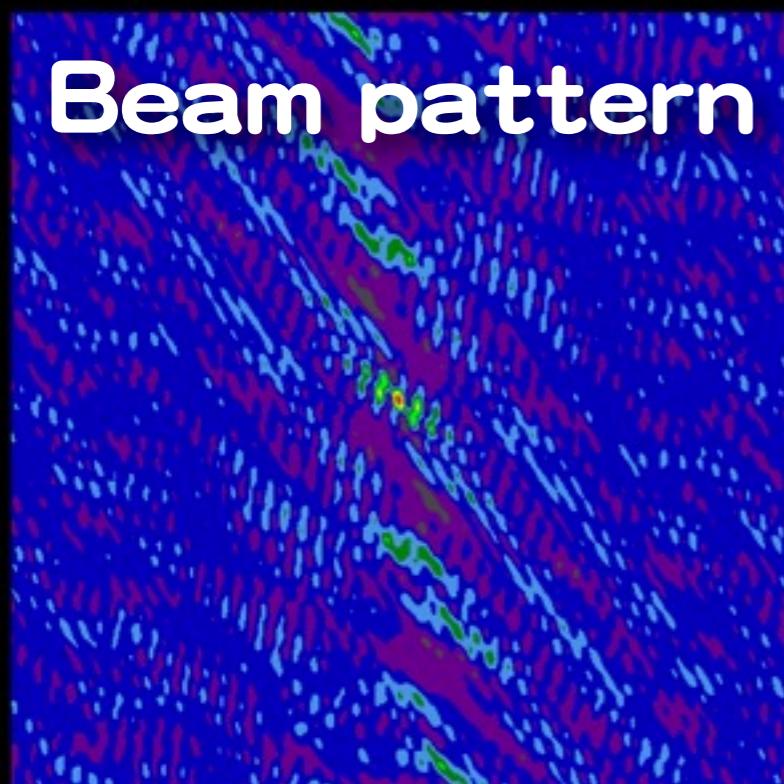
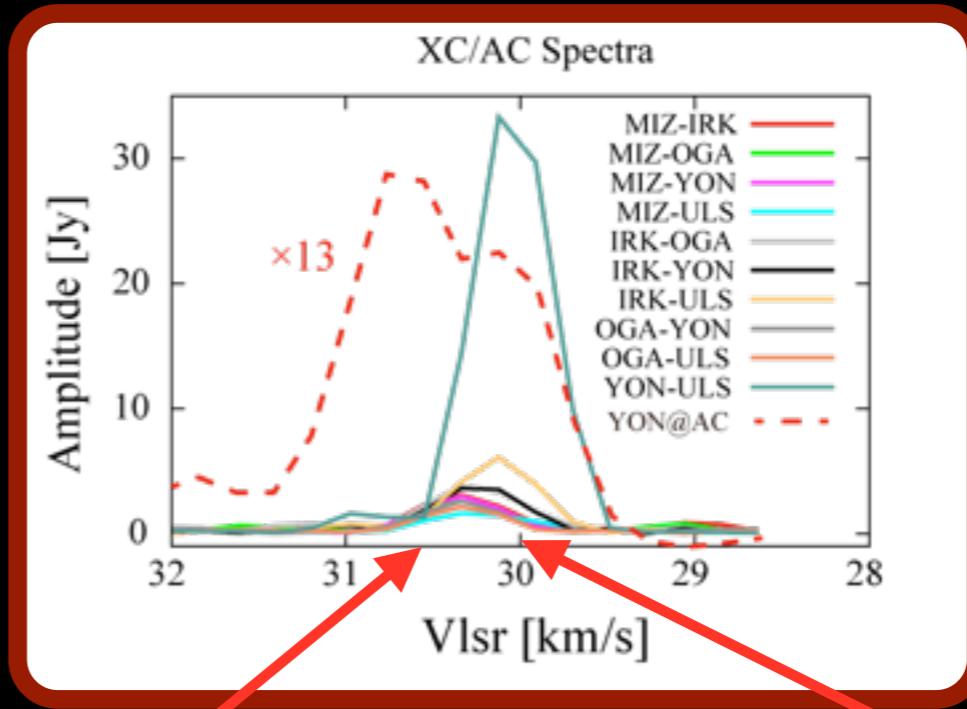
## Spot size estimation



Suggests Compact + Extended component.

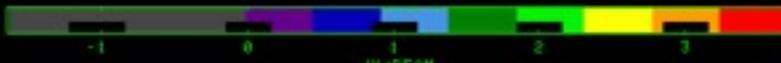
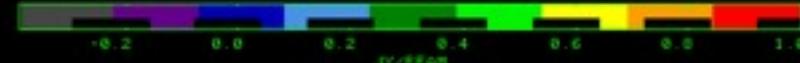
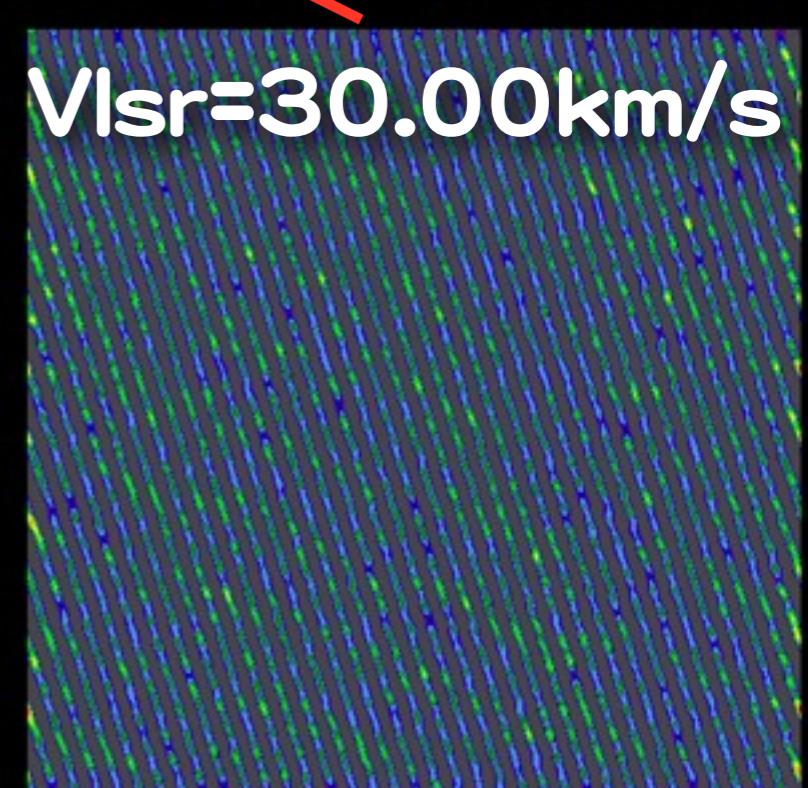
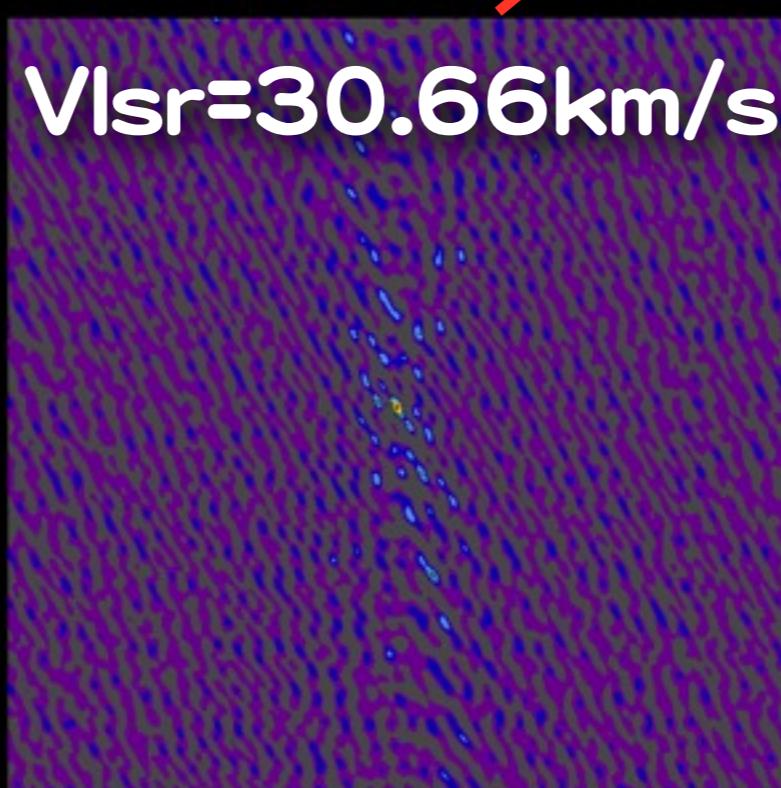
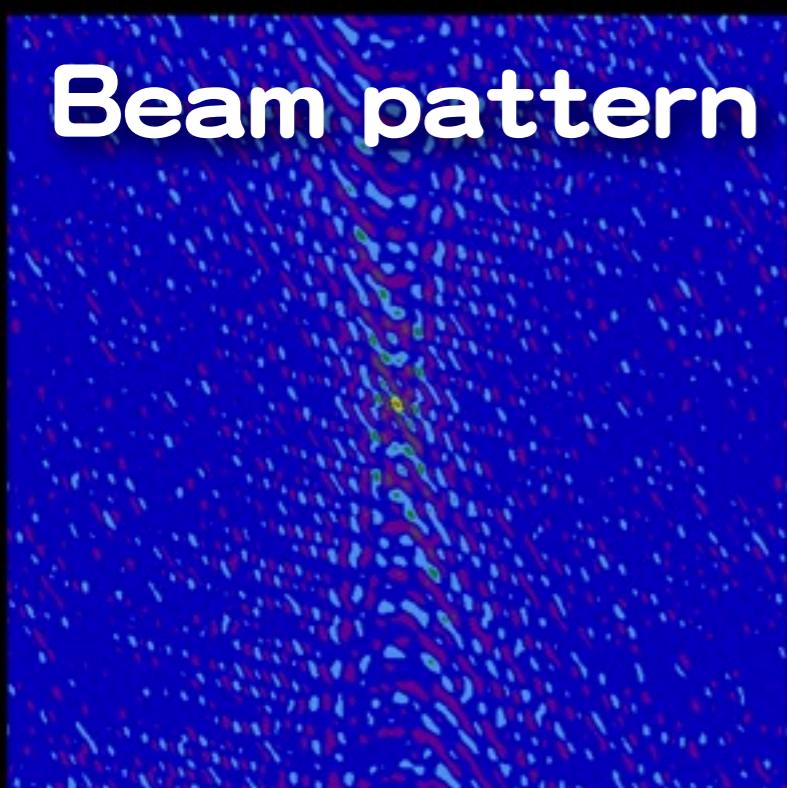
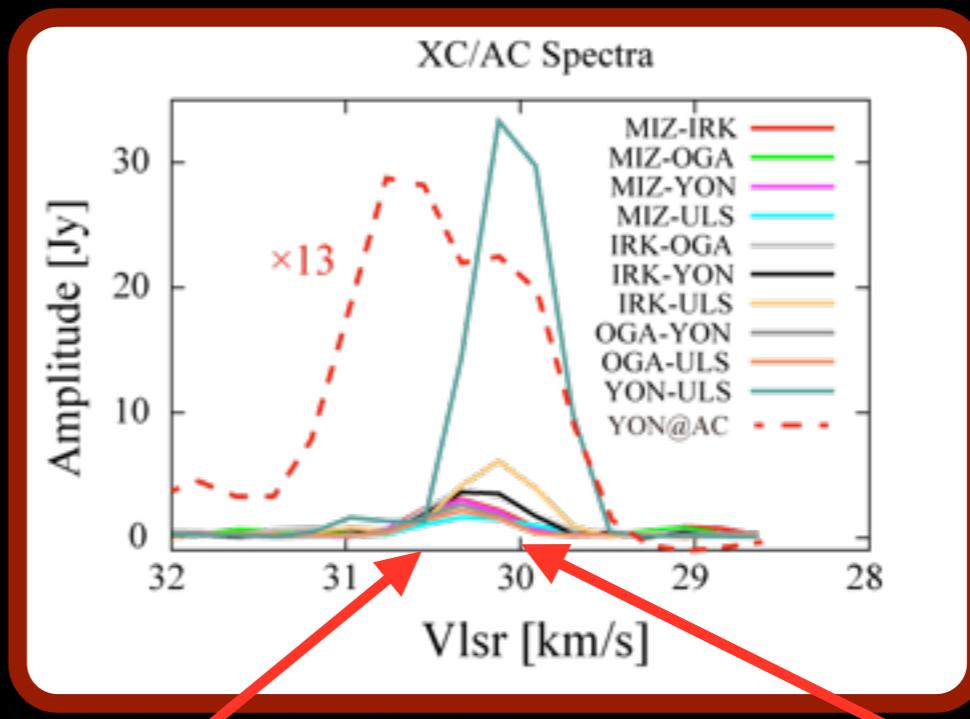
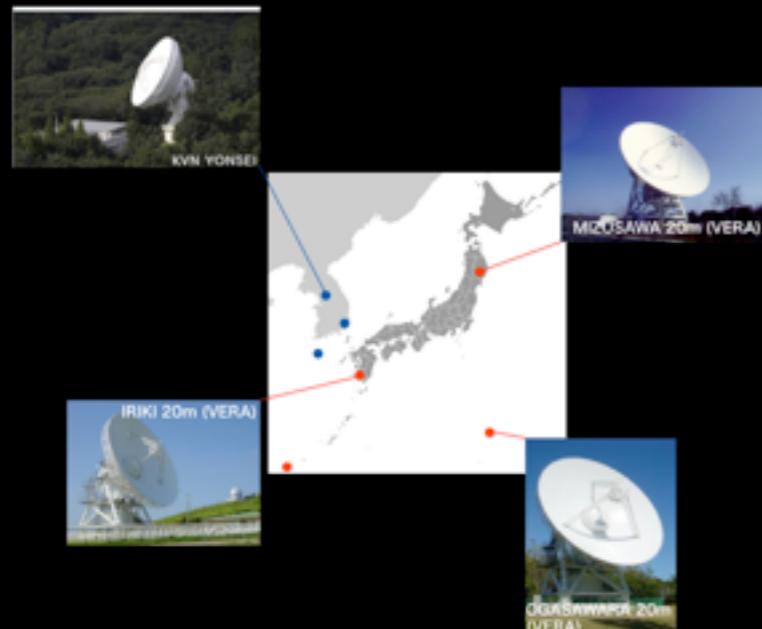
# Results

## The world's first 44GHz methanol VLBI maps



# Results

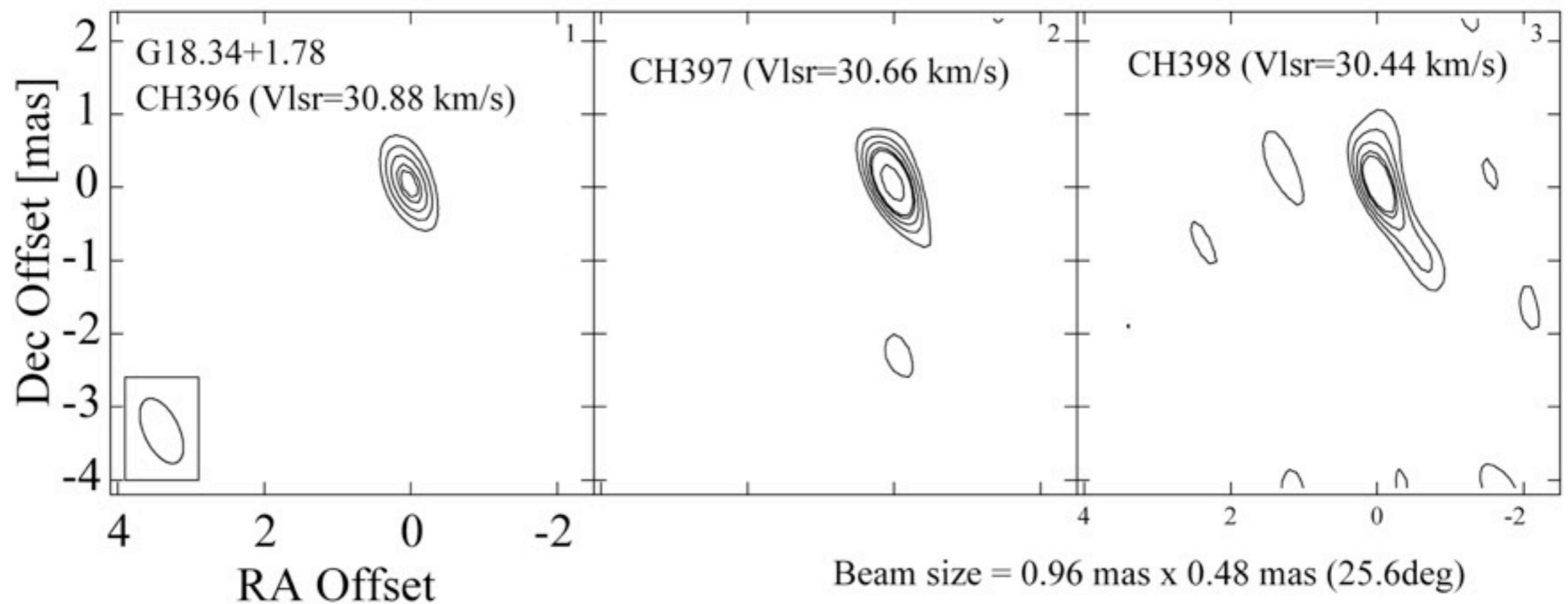
## The world's first 44GHz methanol VLBI maps



# Results

The world's first

## 44GHz methanol VLBI maps



A compact component was detected  
in 3 channels.

# Summary

- We conducted imaging observations of 44GHz methanol maser with KVN+VERA first time.
- KVN+VERA observations produce a complementary effect.  
→ Antenna number, efficiency, baseline length, UV-coverage.
- KVN+VERA array was able to detect 44GHz methanol maser toward G18.34+1.74.
  - First image of 44GHz methanol maser spots.
  - Minimum limit of extended component is 4.4 mas( $\sim 12$  AU)
  - Compact maser spots were imaged.
- Near future, we will get following informations for multiple sources.
  - Distributions, absolute positions, spot sizes, proper motions etc.

→ G44.9-0.39 (r12089c) is already imaged by Hirota-san yesterday.

*Fin.*