



12th VERA user's meeting, Mitaka, Japan, 24-25, Sept. 2014

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VCS(VLBI Correlation Subsystem)

 VCS supports various correlation mode such as W1, C1, C2, C3, C4 and C5.

OCTDISK

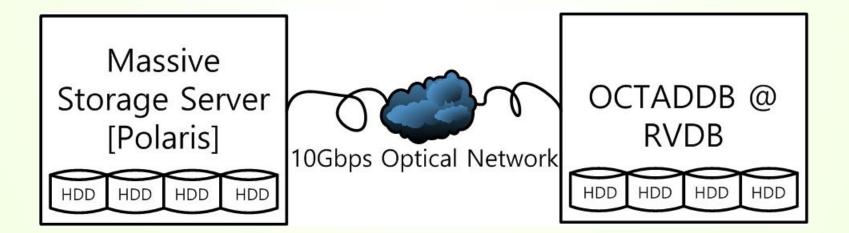
- In RVDB system, 3 OCTADDB system were changed to OCTADISK so as to support VERA(JVN) wideband observation.
- 3 OCTADISKs arrived at KJCC, and playing back and corr was started.



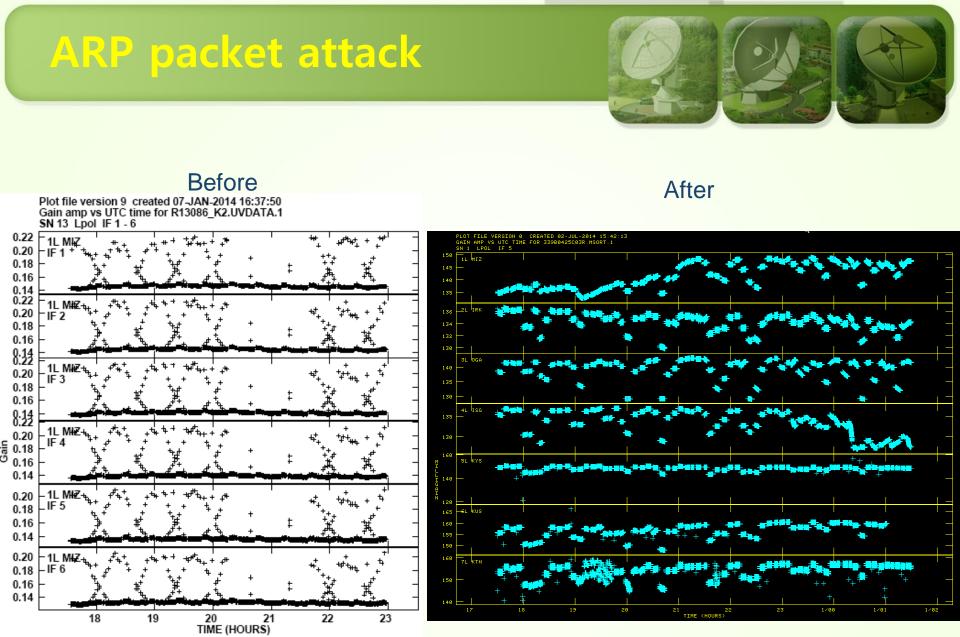


VDIF Copy software : Between massive storage and OCTADDB, VSI obs data is directly possible to copy.

octa_control_rec/octa_control_play

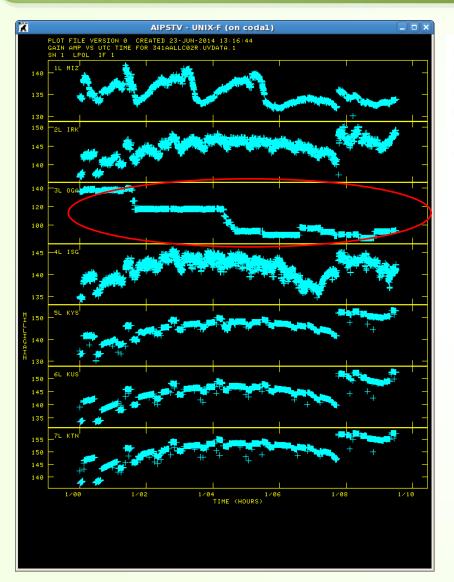


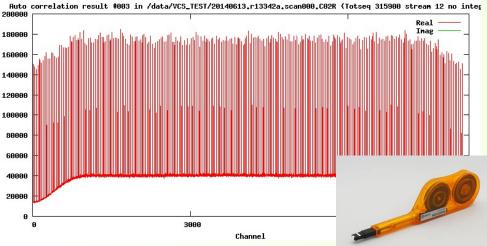






Step-like Gain Amp after ACCOR





- In optical connection between AU and CAB, dust is the reason.
- After cleaning dust, the above problem was disappeared.



KJCC correlation status



KJCC web

<u>http://kjcc.kasi.re.kr/doc/corr-status.html</u>

FITS delivery

 DQA will announce to PI by e-mail to download the FITS file with its link.

***FITS regulations**

- <u>http://kjcc.kasi.re.kr/doc/fits-dist.html</u>
- KJCC would like to receive PI's response within 2 weeks after FITS file downloading. At least if there is no response within 4 weeks, KJCC determined that PI is regarded as his response as "Success" for FITS file(2013.09)

Correlation Status

Season	Observation		Corr Finishe	d Remain C	orr FITS release
2014B	6(in Sep)		-	-	
2014A	57 (1 cancelled, 1 test)		24	31	13
2013B	27		27	0	27
2014A	Total	Disk pa	ck used	Tape used	Remark
VERA	55	29		26	
KVN		55			

2013B, 2014A season

- 84 obs.
- 51 corr. finished



Popularity by operator

- 1. Normal observation with strong fringe finder
- 2. Frequency switching

3. Fast switching

- Fast switching obs is OK, but start/end time in obs log had been superposed because of the different telescope slewing speed.
- If you want to observe this mode, please consider start/end time between scans.
- 4. A/B-beam, Multi-Frequency Obs

∻5. Sub-array

 Observation is just 1 time, but correlation should be performed many times. It takes many hours to correlate them.



KJCC Request(1)



Short cover report of your observation introduction

Discuss it DM to improve operation effectively

How to correlate your observation

- Our operator had some experience recently, but sometimes PI conducted new observation without any introduction. In that case, correlation time is more needed by understanding the observation why it was conducted that way.
- Difference between VEX and Obs log
 - Sometimes we found that VEX and log have different information such as frequency setting or LSB/USB or RHCP/LHCP etc.





Support various obs mode

- W1 : 512MHz BW 1stream
- C1: 256 MHz BW 1stream
- C2: 128 MHz BW 2streams
- C3: 64 MHz BW 4streams
- C4: 32 MHz BW 8streams
- C5 : 16 MHz BW 16streams : Normal
- ☆ → KJCC suggest to PI to do long obs by using W1~C4 mode, which will be used to do your research and to give confirmation of KJCC performance with risk-shared.

How to correlate rapidly



- Stable operation of VCS
- OCTADISK pack
- Simple observation mode w/ strong fringe finder
 - No Fast switching
 - No sub-array
- VEX and log consistence
- Short correlation request
- Quick response after analyzing FITS within 1 month



OCTADISK-pack



Advantage

- Copy work is not needed
- Fast correlation is possible
- PI will quickly receive FITS file

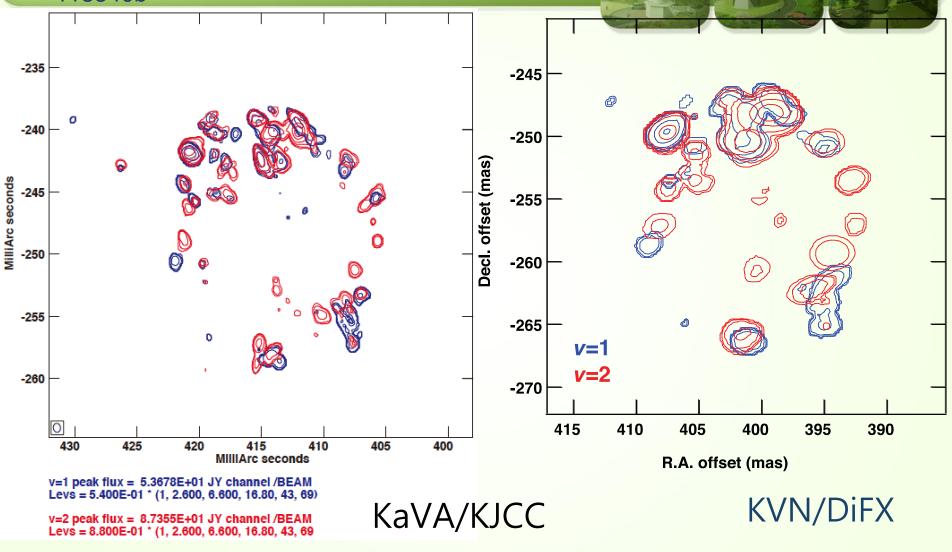
Disadvantage

- Point of view by operator, strong pressure from PI(AOC)
- Busy to correlate
- Fast release of disk-media



SiO masers around BX Cam

r13340b



By Evolved Star Group(H.Imai et al.)

Korea Astronomy & Space

KJCC operation



People in KJCC :

- Roh(Man), Oh, Yeom(HW/SW), Yun(SW), OhCS, Miyazaki(DQA), Kim(DiFX), Jung, Hwang(Op)
- New operator arrived : Hwang Ju-yeon from August





Archive system upgrade is needed Polarization correlation

Test correlation will be also performed soon.

