VERA User's Meeting September 24, 25, 2014

The Tian Ma 65-m telescope in the East Asia VLBI

N. Kawaguchi Shanghai Observatory

Location of the TM65m



I am now working at Sheshan Campus.





The Sheshan Campus



Geometry around Sheshan



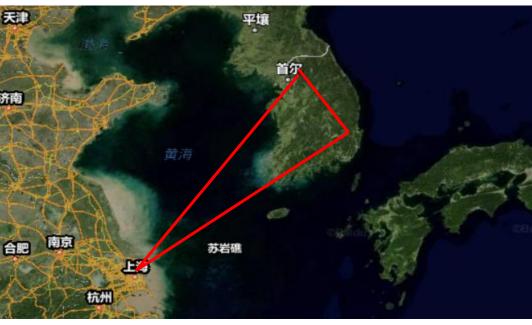
East Asia VLBI Observation on May 27, 2014



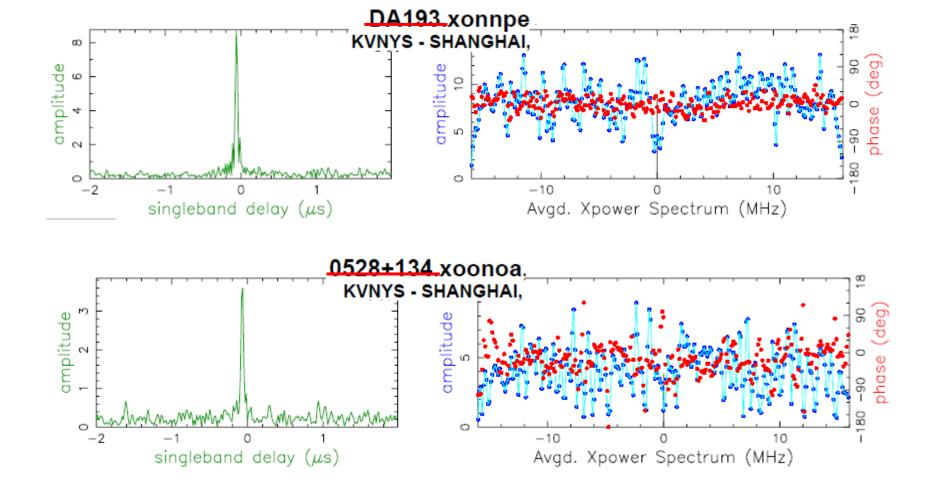
K-band Fringe Test

The Sheshan 25m telescope was participated with the KVN.





K-band fringes detected in Sheshan Correlator



Sensitivities calculated from PCSnr

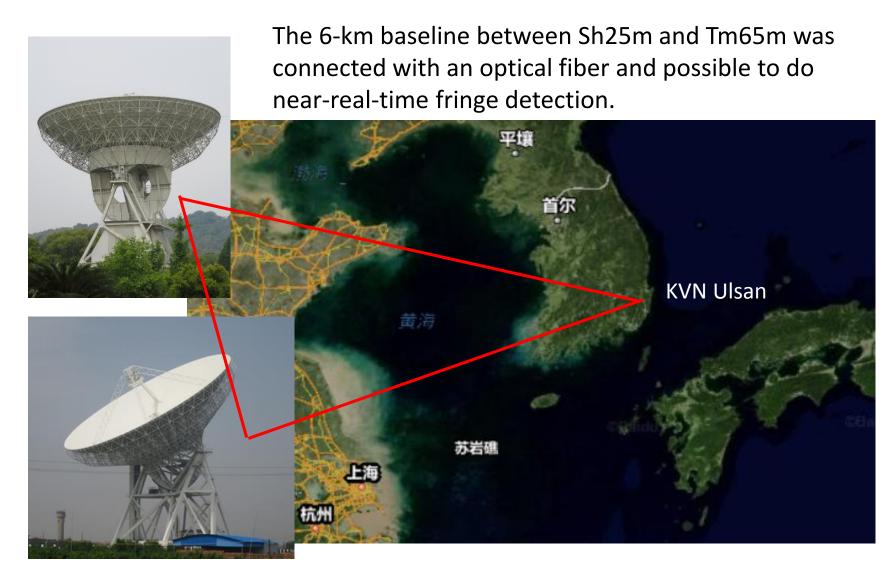
DA193 is referred as 1.75Jy

Stn Name (Stn index**)	KVNUS(i)	KVNYS(k)	Sh25(j)
PCsnr (Observed)	70.2	69.2	28.7
SEFD Observed (Jy)	1372	1392	3356
SEFD Expected (Jy)	1368 (60%) (103K)*	1450 (60%) (114K)*	2700 (50%) (250K)
Stn Name (Stn index)	KVNUS(i)	KVNYS(k)	Sh25(j)

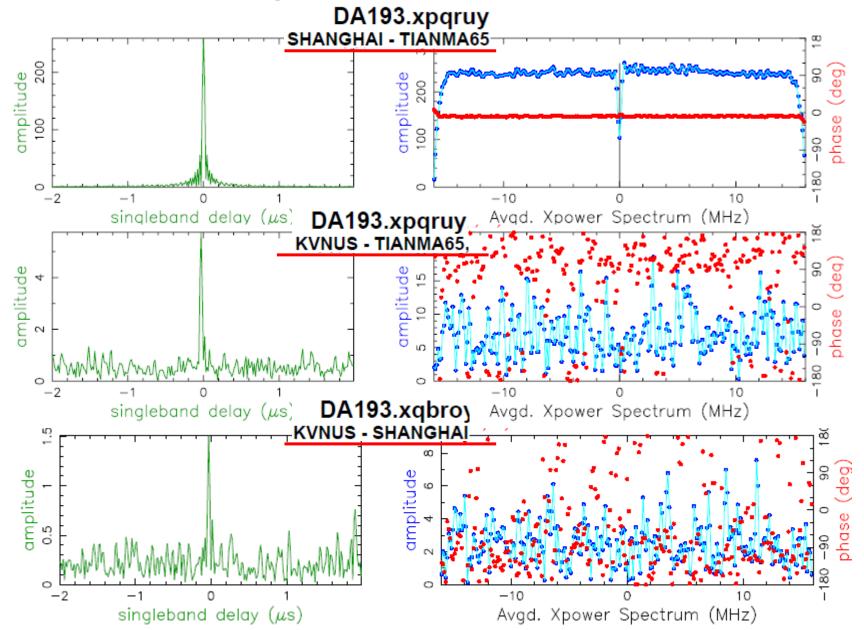
0528+134 is referred as 2.01Jy, 80% resolved

		(103K)*	(114K)*	(250K)
Stn Name (S	tn index)	KVNUS(i)	KVNYS(k)	Sh25(j)
PCsnr (Obser	rved)	45.2	36.3	10.7
SEFD Observ	ed (Jy)	1242	1546	3360
SEFD Expecte	ed (Jy)	1222 (60%) (92K)*	1328 (60%) (100K)*	2700 (50%) (250K)

X-band Fringe Test



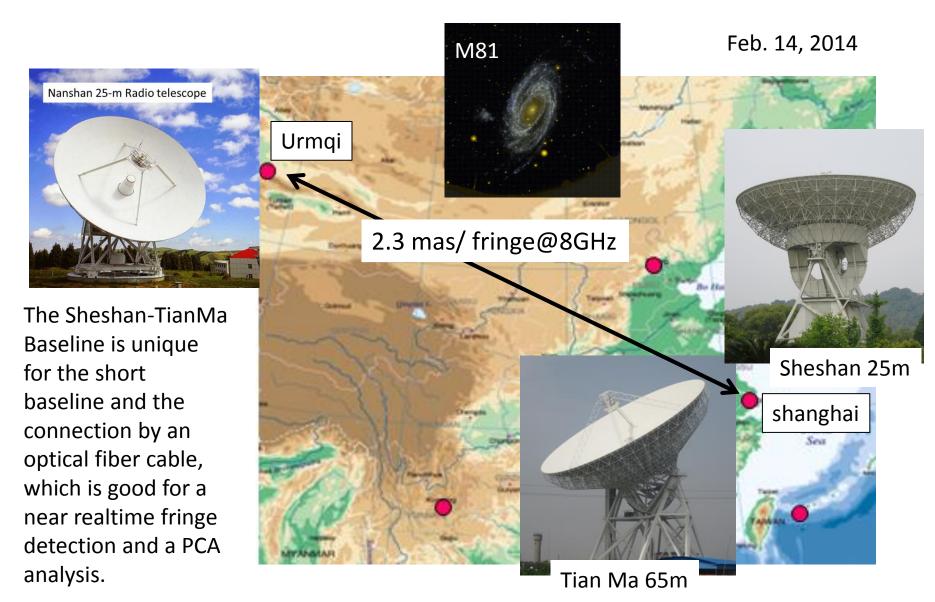
X-band fringes



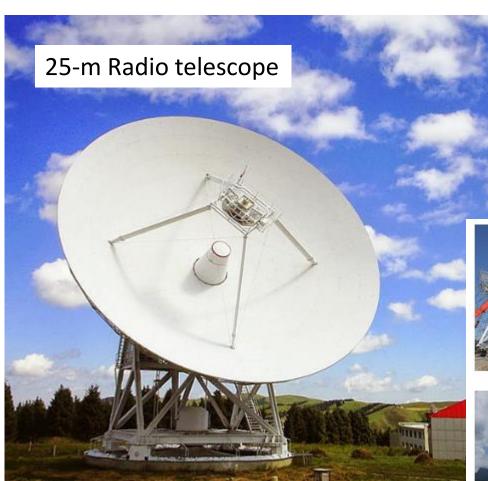
Sensitivities in X-band analyzed on the PCA, Pseudo Closure Amplitude

Station	Code	PCA	SEFD (Jy)	Prev. Result
KVNUS	U	3.322x10 ⁻⁶	2.17x10 ⁶	
Shanghai 25m	S	6.854x10 ⁻³	1050	900
TianMa 65m	Т	9.865x10 ⁻²	73	77
Shanghai 25m – TianMa 65m	ST		277	273 (DA193) 254 (0518+134)

M81 with three CVN stations



Nanshan 25-m, Urmqi



X=228310.3726m

Y=4631922.7697m

Z=4367064.0425m

Under reforming to 26-m diameter

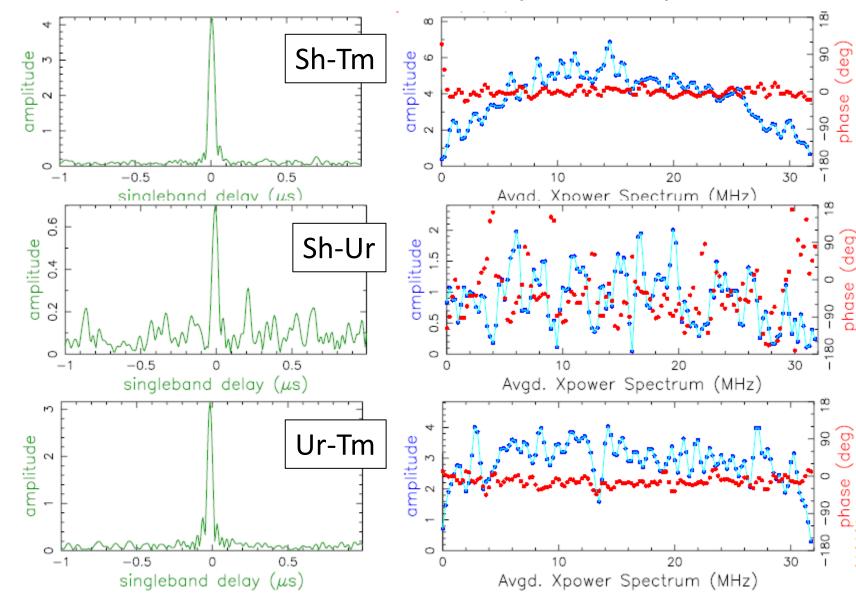




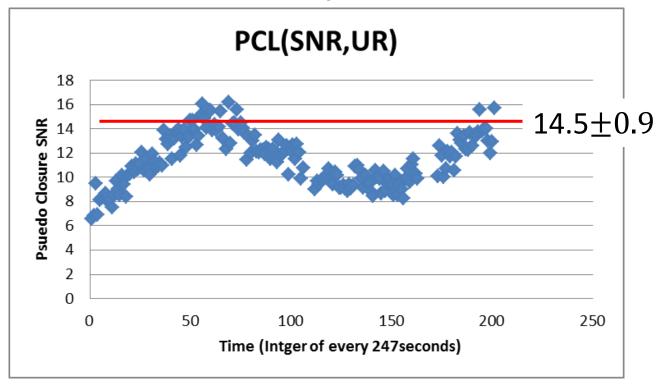




M81 Observation on Sh25, Tm65, Ur25



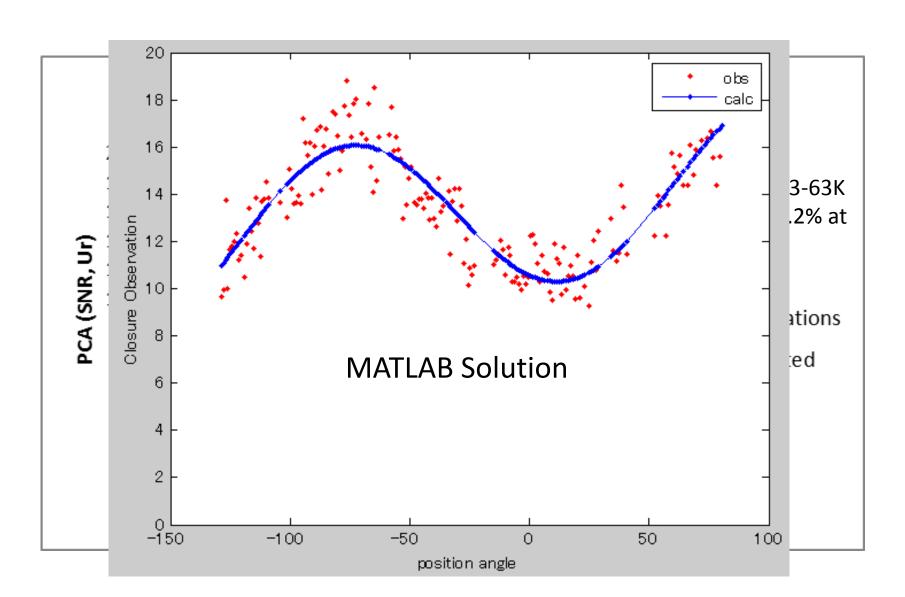
Pseudo Closure Analysis and Sensitivity of Ur25m



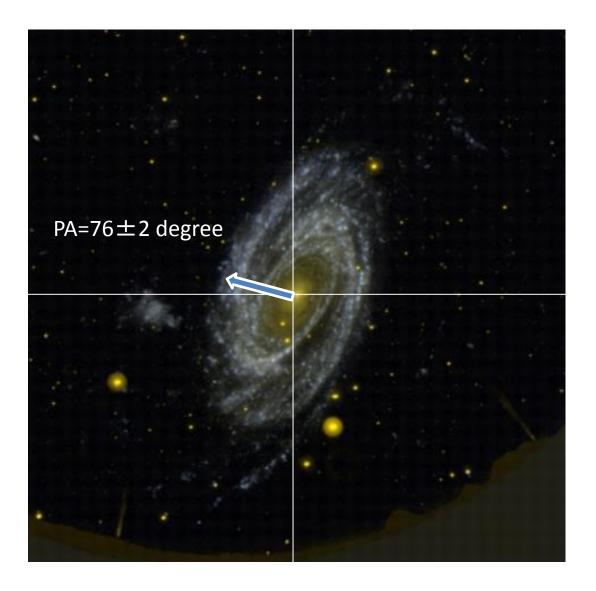
$$SEFD(Ur) = \frac{q \cdot S(M81)\sqrt{2BT}}{PCA(SNR, Ur)} = \frac{0.86 \times 0.17 \times (1.07 \times 10^5)}{(14.5 \pm 0.9)} = \mathbf{1078 \pm 70} [Jy]$$

(The SEFD(Sh) is 1050 Jy obtained in the EAVN fringe test in May, 2014)

Two point source model



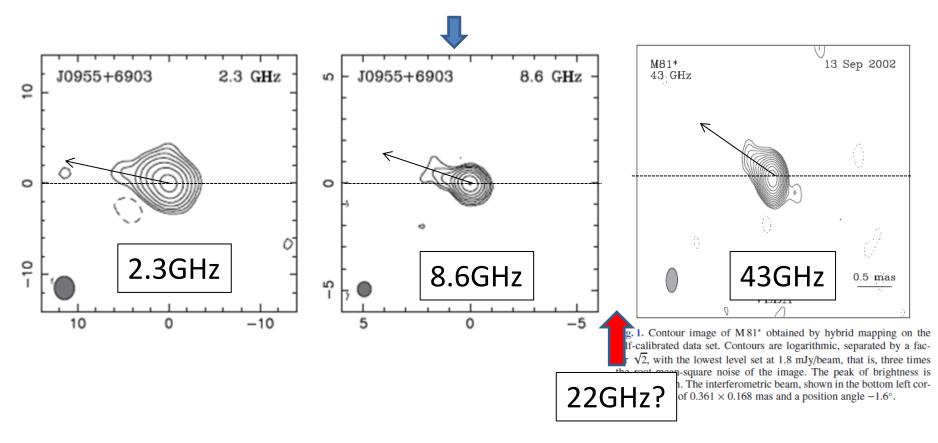
Jet Direction



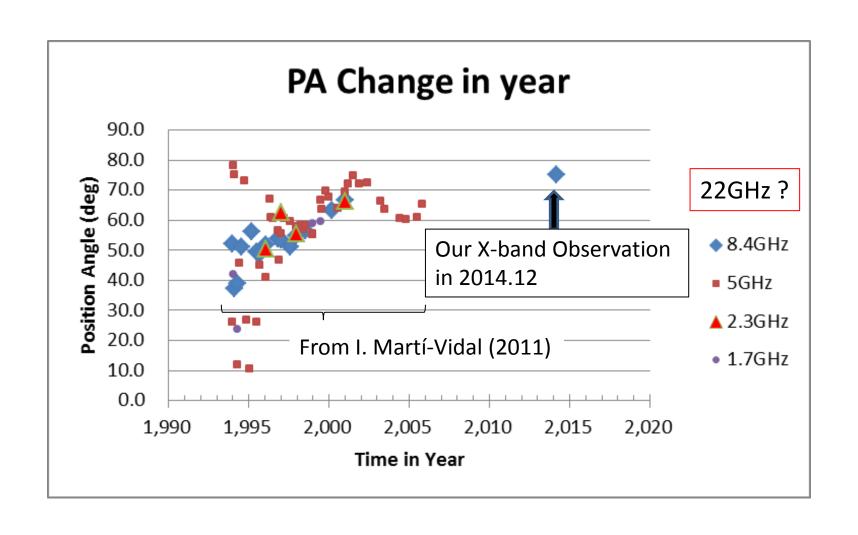
- We can give a twopoint-component
 model to the change
 of PCA with the
 correlation coefficient
 of 99% as seen in the
 previous slide.
- We got peak PCA at the position angle of 75 degree.
- The Jet component is located at almost perpendicular to the galactic plane.

Is the position angle changing with frequency?

Our observation confirms this in X-band.



Jet Precession?

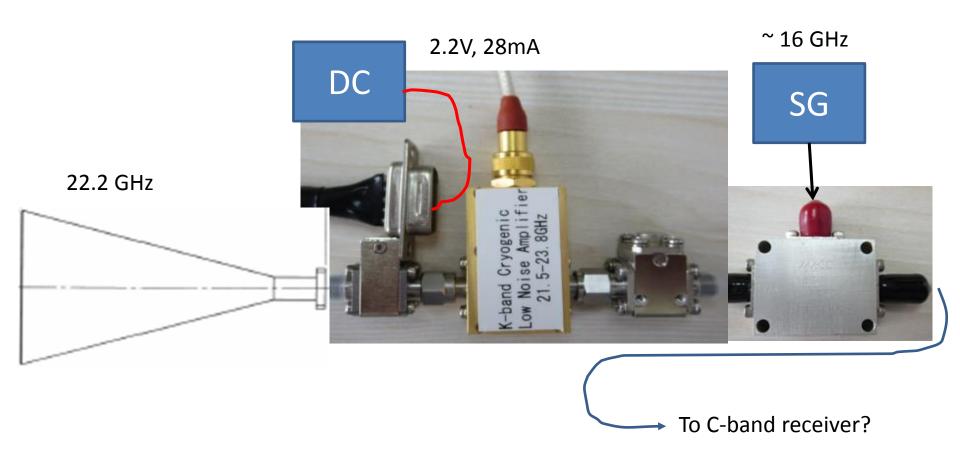


Holography test on a 65-m main reflector

TOWARD 22GHZ

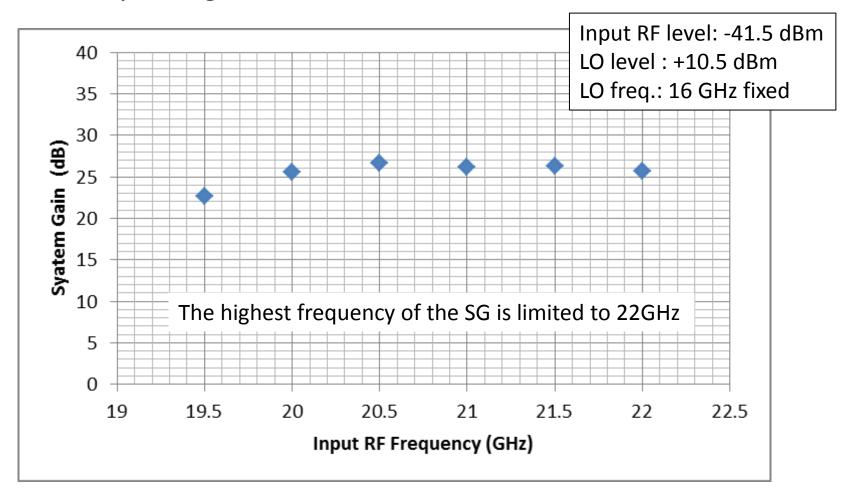
22-GHz test receiver set up

A test 22-GHz receiver only for the use in the holographic surface measurement.



Frequency Response

A system gain of the LNA and the Mixer is 25.5 dB.



Test Staffs



Chen Ying-san (Left) and Liang Zhan Gang-san (Right)

Concluding Remarks

- The Tian Ma 65-m telescope is a powerful telescope which improves the sensitivity of the EAVN much more.
- The 65-m will soon be equipped with a 22-GHz receiver which enables us to use in the major frequency band of the VERA and KVN.
- The 65-m already has a 6-GHz receiver which is quite effective to use in the Methanol observations with JVN.

Sheshan Campus guide, Holography Test etc.

SUPPLEMENTS

Sheshan Campus

Main Building



Sheshan Campus



My Office in Sheshan



Sheshan Campus

Dormitory



Sheshan Campus

Dining Hall



A water channel in the campus



65-m View from Campus

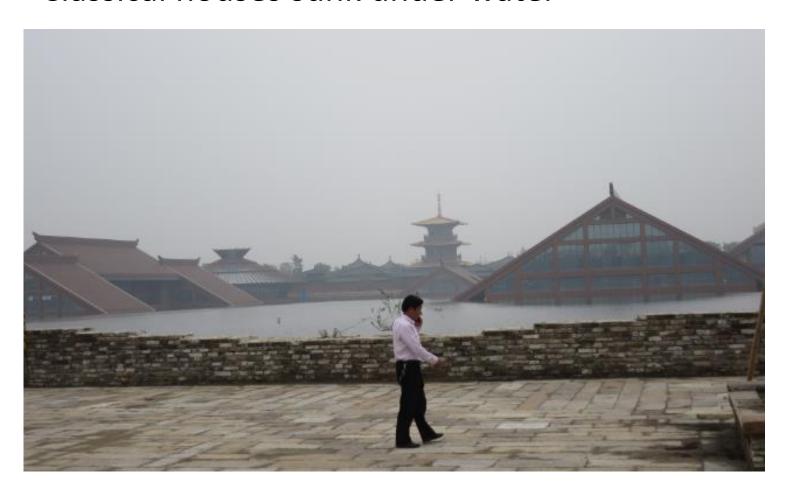


Electric Motor Cycle



Sight Seeing around Sheshan

Classical houses sunk under water



SongJian City







