

# VERA のデータ解析結果の報告

2016 年 10 月 3 日

永山匠、VERA プロジェクトチーム

## 1 解析結果

VERA で観測済みの水メーザー源約 180 天体のうち、未解析のデータと再解析が必要なデータの解析を進めた。35 天体解析し、27 天体で固有運動・年周視差が測定できた。表 1 に解析した天体リストを示す。

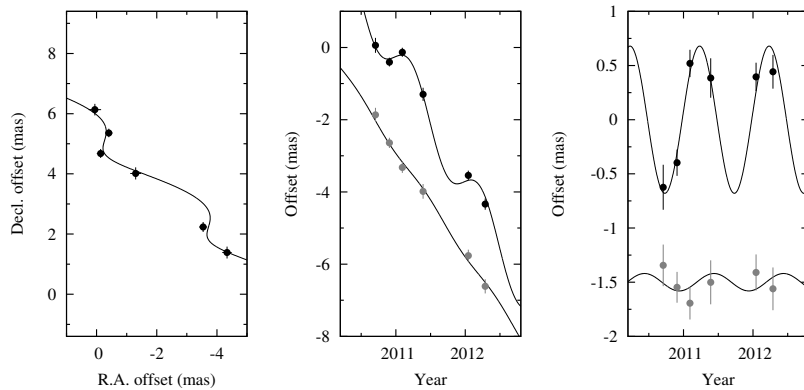
表 1: 天体リスト

No	Name	$l$ (deg)	$b$ (deg)	$\alpha$ (J2000)	$\delta$ (J2000)	$\pi$ (mas)	$\mu_\alpha \cos \delta$ (mas yr <sup>-1</sup> )	$\mu_\delta$ (mas yr <sup>-1</sup> )	$v_{\text{LSR}}$ (km s <sup>-1</sup> )
1	IRAS1814	13.657	-0.599	18h17m24.2446s	-17d22'12.782"	...	...	...	50.6
2	IRAS1815	14.228	-0.509	18h18m12.3118s	-16d49'27.994"	0.693 ± 0.137	-3.36 ± 0.16	-2.88 ± 0.13	9.0
3	02301-00	23.010	-0.411	18h34m40.3191s	-09d00'38.132"	0.304 ± 0.045	-0.39 ± 0.09	-4.61 ± 0.18	79.5
4	G35.03	35.025	0.350	18h54m00.6456s	+02d01'19.393"	0.457 ± 0.025	-1.44 ± 0.03	-4.36 ± 0.07	44.3
5	W48	35.200	-1.736	19h01m45.5423s	+01d13'32.573"	0.394 ± 0.029	-0.50 ± 0.71	-4.11 ± 0.42	41.1 ± 1.2
6	I18517	37.430	1.518	18h54m14.2947s	+04d41'40.559"	0.505 ± 0.059	3.48 ± 0.09	-2.26 ± 0.15	37.6
7	18553+04	37.498	0.530	18h57m53.3876s	+04d18'17.394"	0.077 ± 0.018	-2.75 ± 0.12	-5.52 ± 0.07	10.8 ± 2.2
8	18563+04	37.820	0.413	18h58m53.8800s	+04d32'15.004"	0.082 ± 0.029	-2.82 ± 0.11	-5.44 ± 0.11	17.9 ± 0.7
9	S76E	40.496	2.539	18h56m11.4413s	+07d53'17.608"	0.521 ± 0.024	-0.89 ± 0.34	-2.27 ± 0.56	31.9 ± 1.7
10	I19087+0	43.261	-0.210	19h11m11.7000s	+09d05'06.000"	...	...	...	...
11	I19088+0	43.307	-0.211	19h11m17.2311s	+09d07'32.135"	0.136 ± 0.023	-2.16 ± 0.04	-5.26 ± 0.05	59.2
12	G45.07	45.072	0.133	19h13m22.0696s	+10d50'55.399"	0.235 ± 0.042	-3.14 ± 0.08	-5.44 ± 0.15	63.4
13	IRAS1918	48.606	0.023	19h20m31.1861s	+13d55'25.176"	0.084 ± 0.015	-2.89 ± 0.03	-5.43 ± 0.04	20.7
14	I19446	61.476	0.094	19h46m47.9175s	+25d12'52.698"	0.279 ± 0.055	-0.24 ± 0.23	-6.35 ± 0.16	59.0
15	I2014+35	73.653	0.194	20h16m21.9324s	+35d36'06.102"	0.119 ± 0.010	-2.49 ± 0.06	-4.07 ± 0.08	-75.6 ± 1.0
16	IRAS2023	74.044	-1.710	20h25m07.8013s	+34d50'34.733"	0.611 ± 0.022	-2.83 ± 0.33	-3.28 ± 0.19	6.8 ± 0.4
17	I2014+37	75.296	1.324	20h16m16.0214s	+37d35'44.888"	0.110 ± 0.014	-2.88 ± 0.02	-4.46 ± 0.02	-57.1 ± 1.8
18	WB93	94.262	-0.411	21h32m31.5000s	+51d02'22.000"	...	...	...	...
19	I22566+5	108.754	-0.988	22h58m46.4000s	+58d44'50.000"	...	...	...	...
20	IC1805W	134.279	0.856	02h29m01.9375s	+61d33'30.999"	0.814 ± 0.013	-1.72 ± 0.11	-1.01 ± 0.18	-42.5
21	G135.28	135.277	2.798	02h43m28.5825s	+62d57'08.390"	0.110 ± 0.009	-0.60 ± 0.16	0.16 ± 0.16	-70.0 ± 1.9
22	IRAS0246	136.384	2.267	02h50m08.9000s	+61d59'53.500"	...	...	...	...
23	G137.07	137.068	3.003	02h58m13.1793s	+62d20'32.915"	0.183 ± 0.012	-0.73 ± 0.14	0.01 ± 0.17	-50.1 ± 0.4
24	NGC1579	165.475	-9.054	04h30m27.4008s	+35d09'17.649"	1.695 ± 0.190	2.84 ± 0.20	-8.32 ± 0.22	7.8
25	WB668	172.874	2.269	05h36m52.4670s	+36d10'58.260"	0.582 ± 0.070	-0.05 ± 0.34	-1.69 ± 0.10	-23.7
26	WB673	173.165	2.355	05h38m00.2573s	+35d59'01.103"	...	...	...	-18.5
27	WB763	188.978	0.933	06h09m08.0300s	+21d38'10.700"	...	...	...	...
28	IRAS0646	209.602	1.195	06h49m37.2030s	+03d30'30.474"	0.452 ± 0.051	-0.69 ± 0.07	0.83 ± 0.13	13.8
29	IRAS0650	211.593	1.056	06h52m45.3205s	+01d40'23.048"	0.283 ± 0.108	-0.72 ± 0.15	0.66 ± 0.19	47.3
30	WB886	212.064	-0.740	06h47m13.3344s	+00d26'05.920"	0.110 ± 0.130	-0.58 ± 0.35	0.77 ± 0.53	55.2
31	IRAS0702	224.337	-2.137	07h04m44.3900s	-11d07'12.450"	0.850 ± 0.070	-0.18 ± 0.91	2.33 ± 0.97	14.6 ± 0.8
32	G226.27	226.273	-0.463	07h14m26.9333s	-12d03'58.884"	0.609 ± 0.070	2.22 ± 0.15	2.72 ± 0.59	16.1
33	IRAS0720	229.561	0.152	07h23m01.8407s	-14d41'32.807"	0.080 ± 0.034	-1.38 ± 0.03	0.60 ± 0.08	57.3
34	IRAS0727	235.685	-1.246	07h30m05.2000s	-20d44'04.500"	...	...	...	...
35	IRAS0742	240.316	0.071	07h44m51.9368s	-24d07'42.811"	0.286 ± 0.035	-1.79 ± 0.04	2.01 ± 0.10	66.4

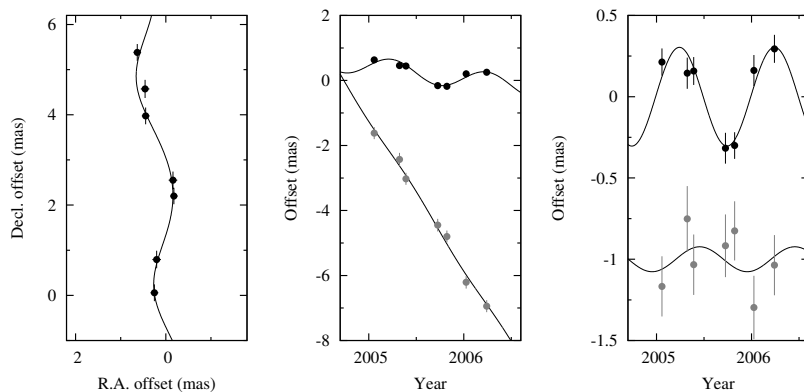
## 2 年周視差

27 天体の固有運動と視差を示す。18553+04, 18563+04, I19088+0, I2014+35, I2014+37, G135.28, G137.07 など 10 kpc 前後の遠方の天体の視差も得られた。

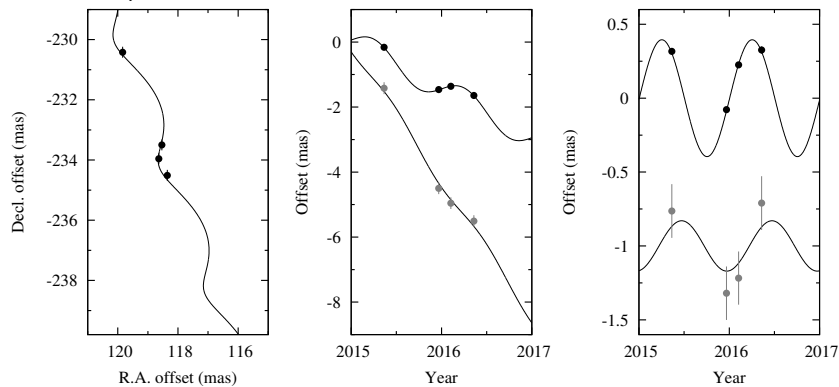
IRAS1815  
 18h18m12.3118s -16d49'27.994"  
 $\pi = 0.680 \pm 0.133$  mas ( $D = 1.47^{+0.36}_{-0.24}$  kpc)  
 $(\mu_x, \mu_y) = (-3.46 \pm 0.14, -2.87 \pm 0.12)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.137, 0.151)$  mas



02301-00  
 18h34m40.3191s -09d00'38.132"  
 $\pi = 0.304 \pm 0.045$  mas ( $D = 3.29^{+0.57}_{-0.42}$  kpc)  
 $(\mu_x, \mu_y) = (-0.39 \pm 0.09, -4.61 \pm 0.18)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.083, 0.187)$  mas



G35.03  
 18h54m00.6456s +02d01'19.393"  
 $\pi = 0.397 \pm 0.023$  mas ( $D = 2.52^{+0.15}_{-0.14}$  kpc)  
 $(\mu_x, \mu_y) = (-1.50 \pm 0.03, -4.17 \pm 0.27)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.014, 0.180)$  mas



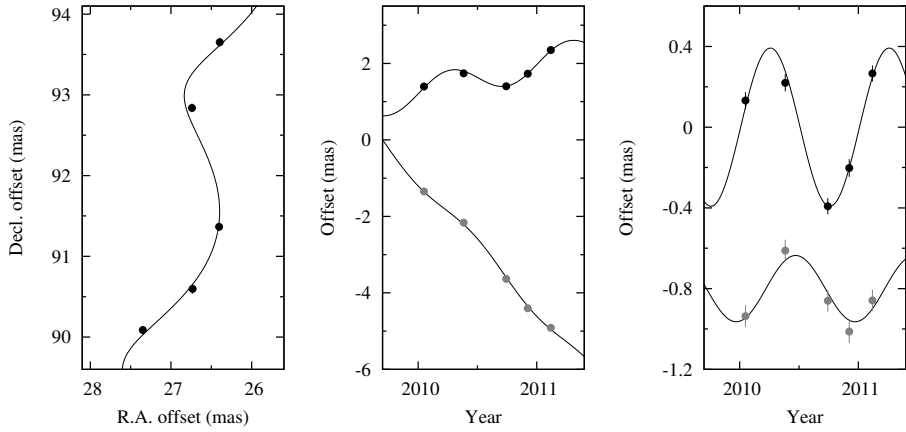
W48

19h01m45.5423s +01d13'32.573"

$\pi = 0.394 \pm 0.029$  mas ( $D = 2.54^{+0.20}_{-0.17}$  kpc)

$(\mu_x, \mu_y) = (0.77 \pm 0.05, -3.41 \pm 0.07)$  mas/yr

$(\sigma_x, \sigma_y) = (0.040, 0.054)$  mas



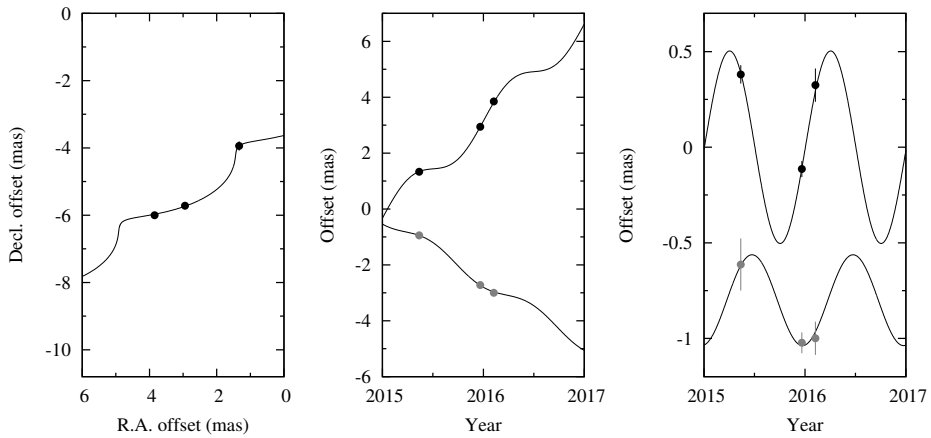
I18517

18h54m14.2947s +04d41'40.559"

$\pi = 0.505 \pm 0.059$  mas ( $D = 1.98^{+0.26}_{-0.21}$  kpc)

$(\mu_x, \mu_y) = (3.48 \pm 0.09, -2.26 \pm 0.15)$  mas/yr

$(\sigma_x, \sigma_y) = (0.030, 0.040)$  mas



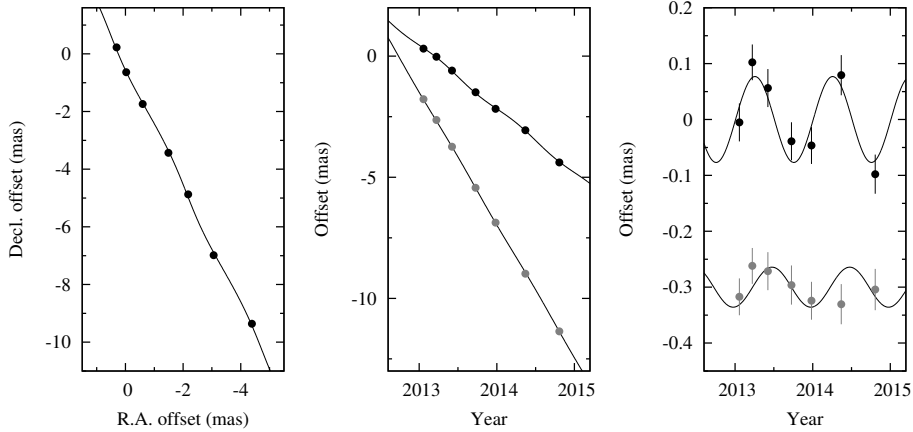
18553+04

18h57m53.3876s +04d18'17.394"

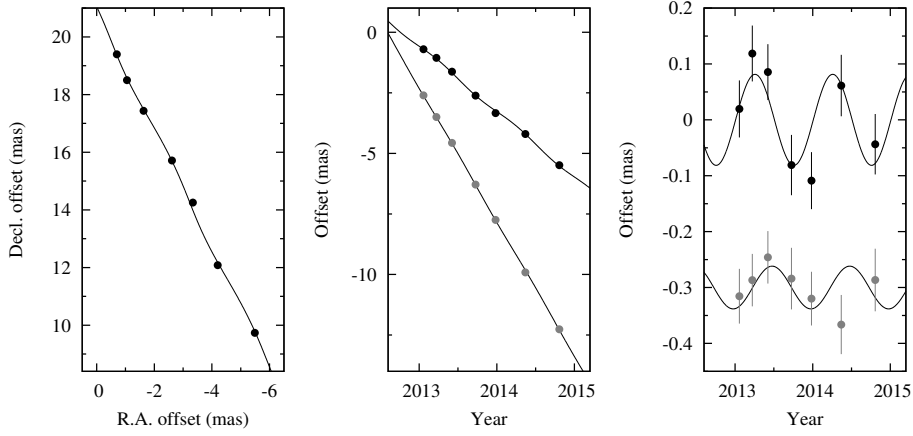
$\pi = 0.077 \pm 0.018$  mas ( $D = 13.0^{+4.0}_{-2.5}$  kpc)

$(\mu_x, \mu_y) = (-2.63 \pm 0.03, -5.49 \pm 0.02)$  mas/yr

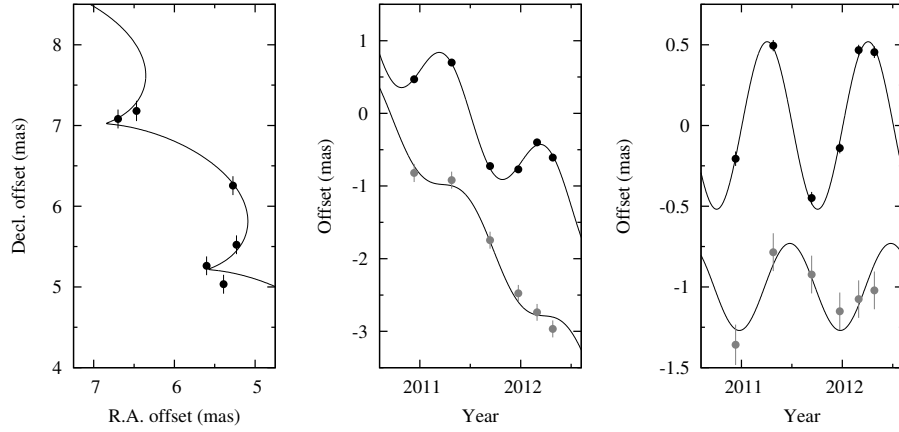
$(\sigma_x, \sigma_y) = (0.033, 0.033)$  mas



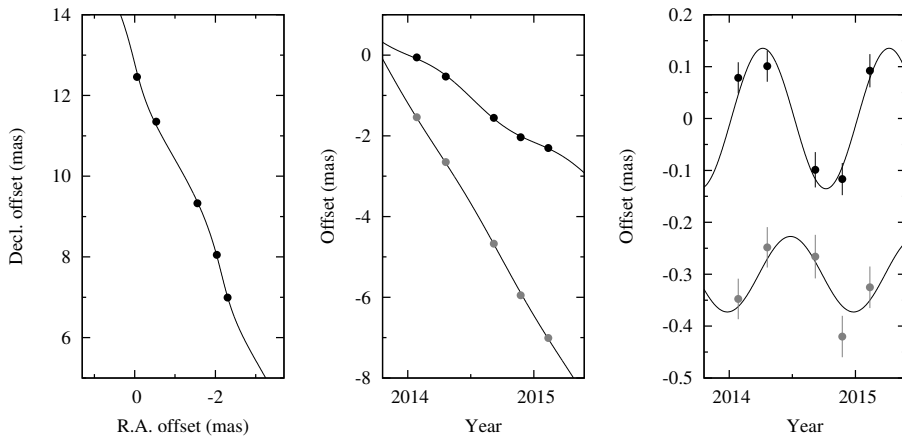
18563+04  
 18h58m53.8800s +04d32'15.004"  
 $\pi = 0.082 \pm 0.029$  mas ( $D = 12.2^{+6.7}_{-3.2}$  kpc)  
 $(\mu_x, \mu_y) = (-2.70 \pm 0.04, -5.54 \pm 0.03)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.053, 0.048)$  mas



S76E  
 18h56m11.4413s +07d53'17.608"  
 $\pi = 0.521 \pm 0.024$  mas ( $D = 1.92^{+0.09}_{-0.08}$  kpc)  
 $(\mu_x, \mu_y) = (-1.27 \pm 0.04, -1.81 \pm 0.11)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.035, 0.120)$  mas



IRAS19088+0902  
 19h11m17.2311s +09d07'32.135"  
 $\pi = 0.136 \pm 0.023$  mas ( $D = 7.35^{+1.50}_{-1.06}$  kpc)  
 $(\mu_x, \mu_y) = (-2.16 \pm 0.04, -5.26 \pm 0.05)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.032, 0.040)$  mas



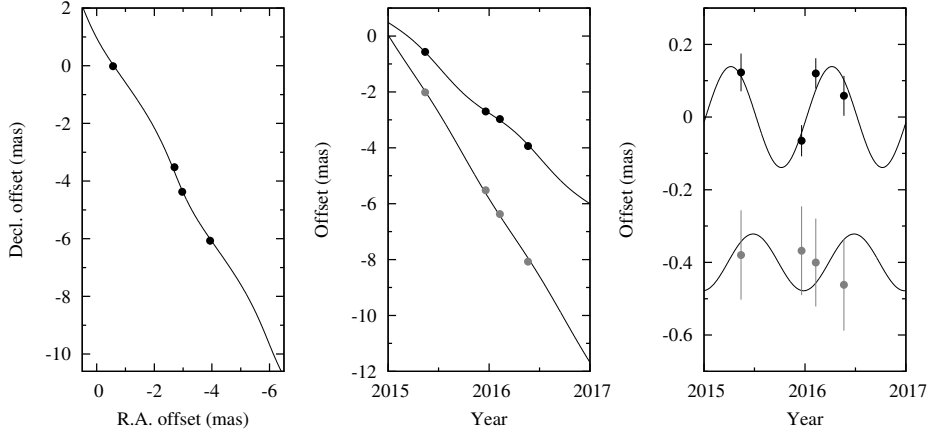
G45.07

19h13m22.0696s +10d50'55.399"

$\pi = 0.140 \pm 0.059$  mas ( $D = 7.1^{+5.2}_{-2.1}$  kpc)

$(\mu_x, \mu_y) = (-3.25 \pm 0.08, -5.86 \pm 0.20)$  mas/yr

$(\sigma_x, \sigma_y) = (0.048, 0.127)$  mas



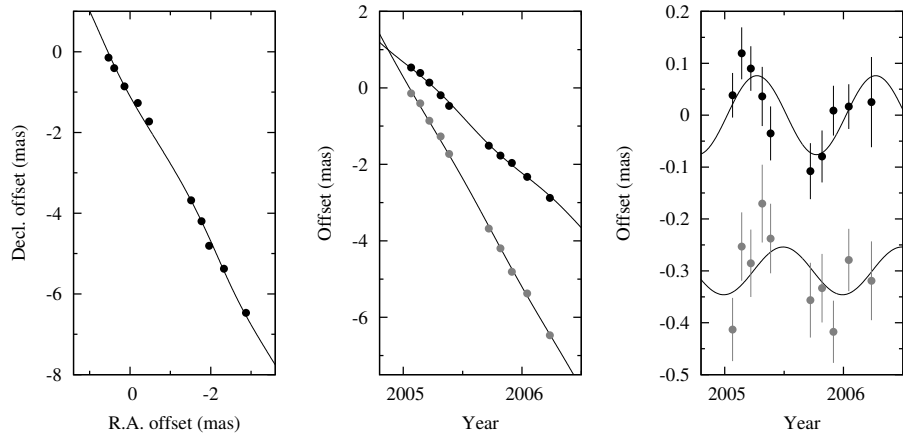
IRAS1918 (TXS1920)

19h20m31.1861s +13d55'25.176"

$\pi = 0.077 \pm 0.026$  mas ( $D = 13.0^{+6.6}_{-3.3}$  kpc)

$(\mu_x, \mu_y) = (-2.90 \pm 0.05, -5.49 \pm 0.06)$  mas/yr

$(\sigma_x, \sigma_y) = (0.052, 0.069)$  mas



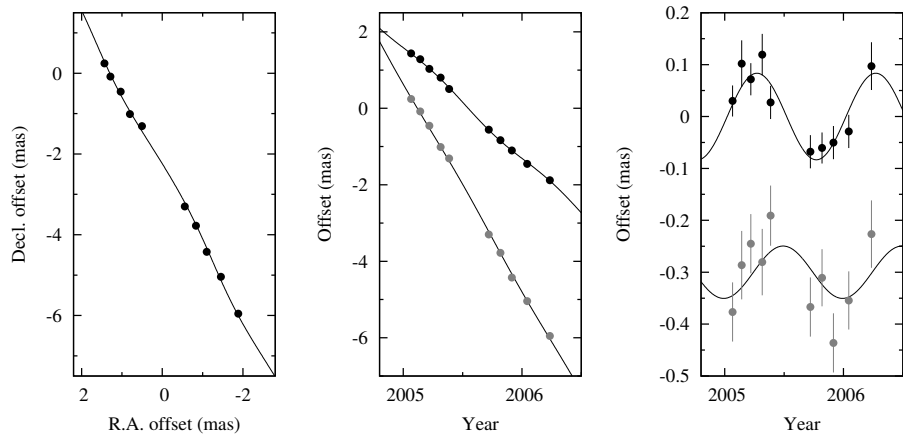
IRAS1918 (TXS1922)

19h20m31.1861s +13d55'25.176"

$\pi = 0.084 \pm 0.015$  mas ( $D = 11.9^{+2.6}_{-1.8}$  kpc)

$(\mu_x, \mu_y) = (-2.89 \pm 0.03, -5.43 \pm 0.04)$  mas/yr

$(\sigma_x, \sigma_y) = (0.030, 0.062)$  mas



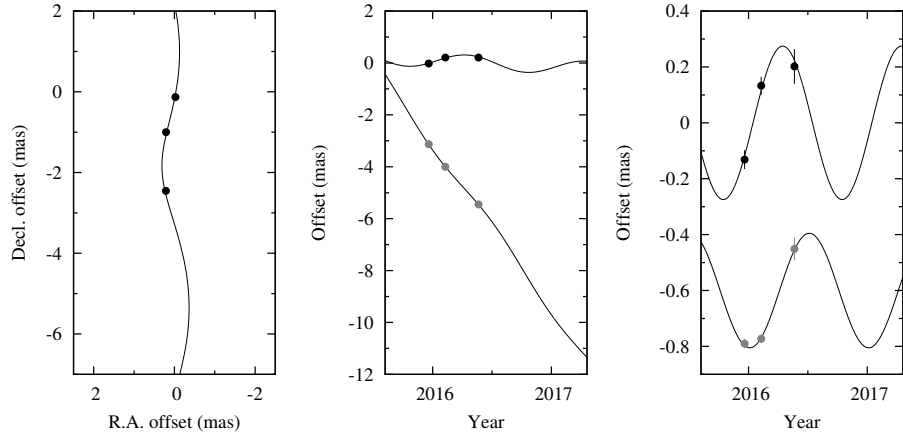
I19446

19h46m47.9175s +25d12'52.698"

$\pi = 0.279 \pm 0.055$  mas ( $D = 3.6^{+0.9}_{-0.6}$  kpc)

$(\mu_x, \mu_y) = (-0.24 \pm 0.23, -6.35 \pm 0.16)$  mas/yr

$(\sigma_x, \sigma_y) = (0.031, 0.009)$  mas



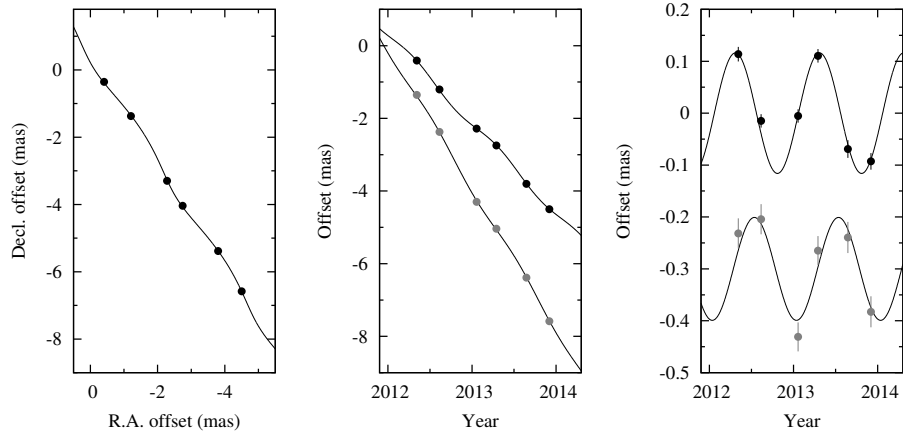
I2014+35

20h16m21.9324s +35d36'06.102"

$\pi = 0.119 \pm 0.010$  mas ( $D = 8.40^{+0.77}_{-0.65}$  kpc)

$(\mu_x, \mu_y) = (-2.46 \pm 0.01, -3.85 \pm 0.02)$  mas/yr

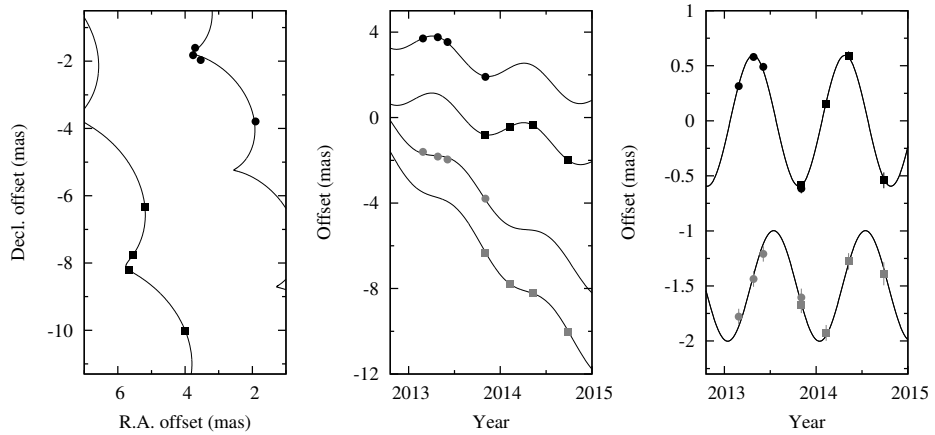
$(\sigma_x, \sigma_y) = (0.013, 0.029)$  mas



IRAS2023

20h25m07.8013s +34d50'34.733"

$\pi = 0.611 \pm 0.022$  mas ( $D = 1.64 \pm 0.06$  kpc)



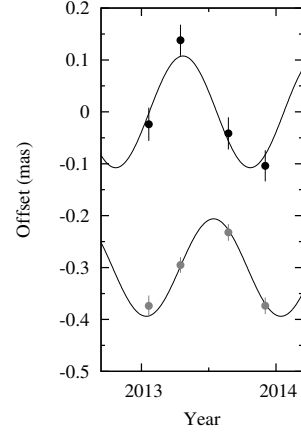
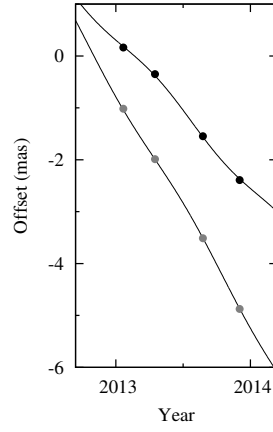
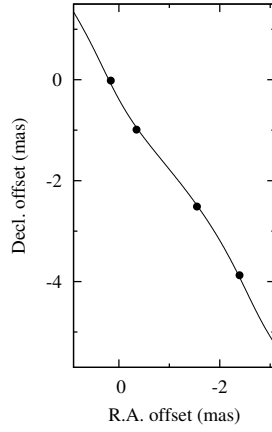
I2014+37

20h16m16.0214s +37d35'44.888"

$\pi = 0.110 \pm 0.014$  mas ( $D = 9.09^{+1.33}_{-1.03}$  kpc)

$(\mu_x, \mu_y) = (-2.86 \pm 0.05, -4.46 \pm 0.03)$  mas/yr

$(\sigma_x, \sigma_y) = (0.032, 0.015)$  mas



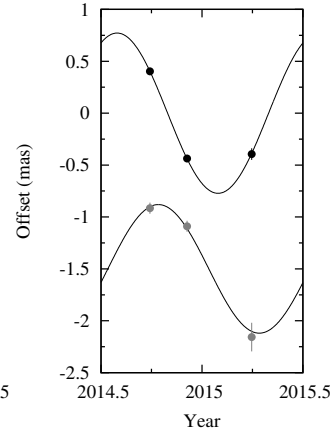
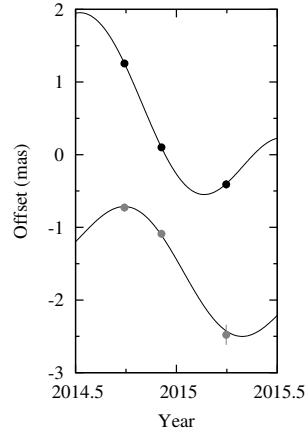
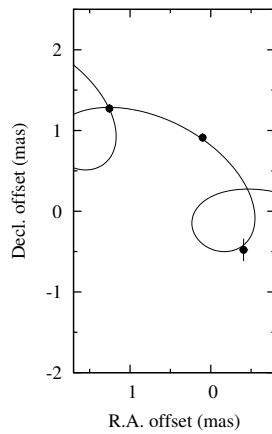
IC1805W

02h29m01.9375s +61d33'30.999"

$\pi = 0.814 \pm 0.013$  mas ( $D = 1.23^{+0.02}_{-0.02}$  kpc)

$(\mu_x, \mu_y) = (-1.72 \pm 0.11, -1.01 \pm 0.18)$  mas/yr

$(\sigma_x, \sigma_y) = (0.007, 0.060)$  mas



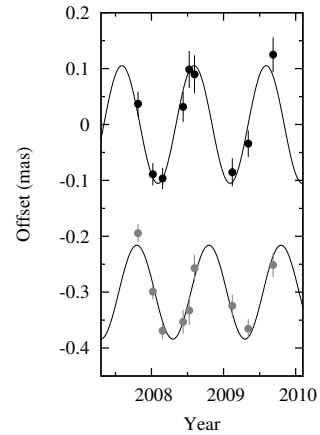
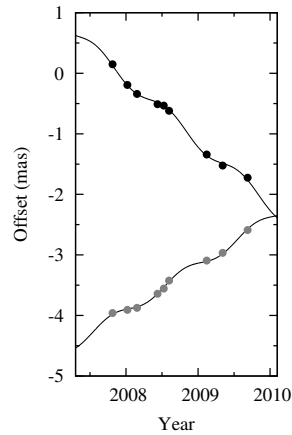
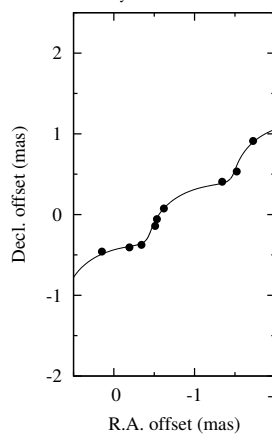
G135.28

02h43m28.5825s +62d57'08.390"

$\pi = 0.110 \pm 0.009$  mas ( $D = 9.09^{+0.81}_{-0.69}$  kpc)

$(\mu_x, \mu_y) = (-1.05 \pm 0.01, 0.76 \pm 0.01)$  mas/yr

$(\sigma_x, \sigma_y) = (0.026, 0.018)$  mas



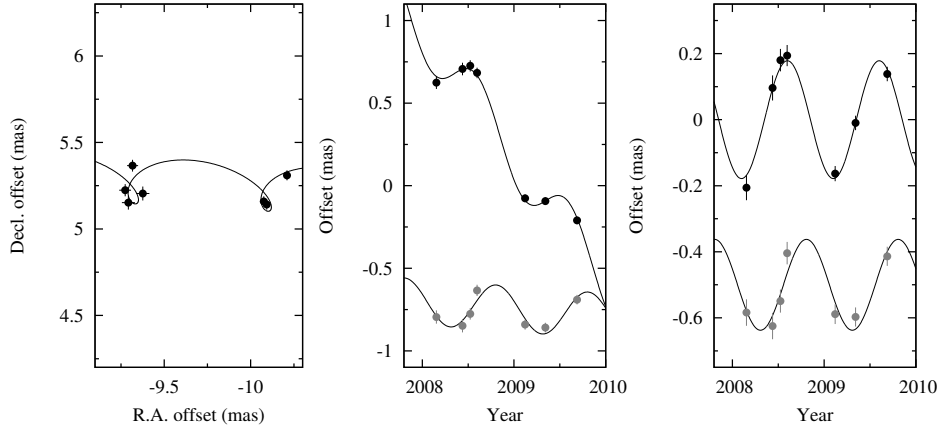
G137.07

02h58m13.1793s +62d20'32.915"

$\pi = 0.183 \pm 0.012$  mas ( $D = 5.45^{+0.38}_{-0.34}$  kpc)

$(\mu_x, \mu_y) = (-0.77 \pm 0.02, -0.04 \pm 0.02)$  mas/yr

$(\sigma_x, \sigma_y) = (0.022, 0.039)$  mas



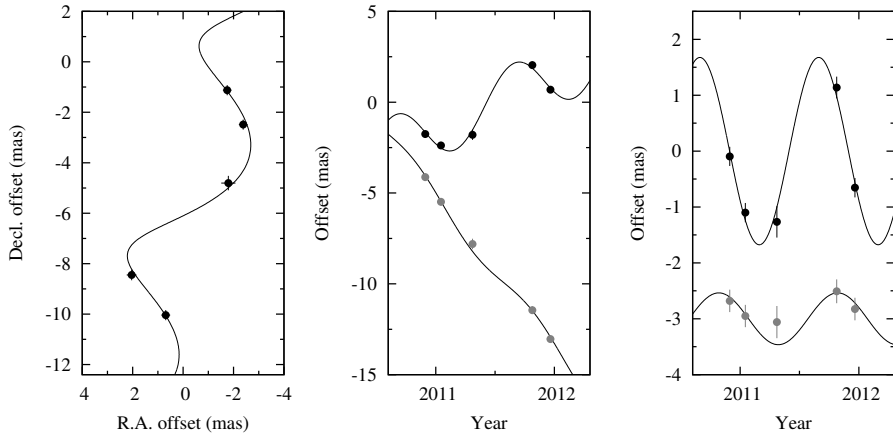
NGC1579

04h30m27.4008s +35d09'17.649"

$\pi = 1.695 \pm 0.190$  mas ( $D = 590^{+74}_{-59}$  pc)

$(\mu_x, \mu_y) = (2.84 \pm 0.20, -8.32 \pm 0.22)$  mas/yr

$(\sigma_x, \sigma_y) = (0.208, 0.245)$  mas



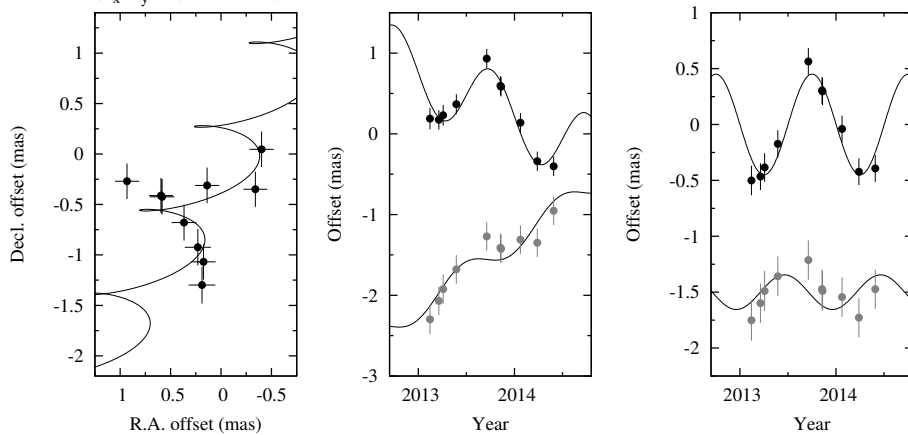
IRAS06469+0333

06h49m37.2030s +03d30'30.474"

$\pi = 0.452 \pm 0.051$  mas ( $D = 2.21^{+0.28}_{-0.22}$  kpc)

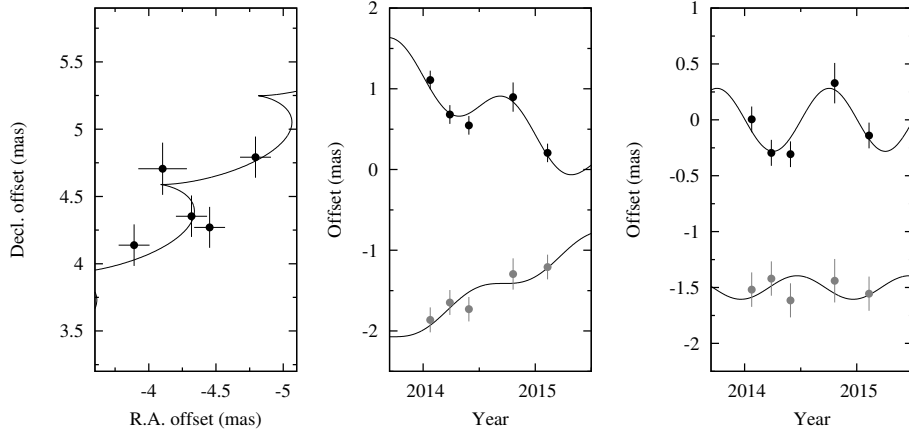
$(\mu_x, \mu_y) = (-0.69 \pm 0.07, 0.83 \pm 0.13)$  mas/yr

$(\sigma_x, \sigma_y) = (0.119, 0.174)$  mas

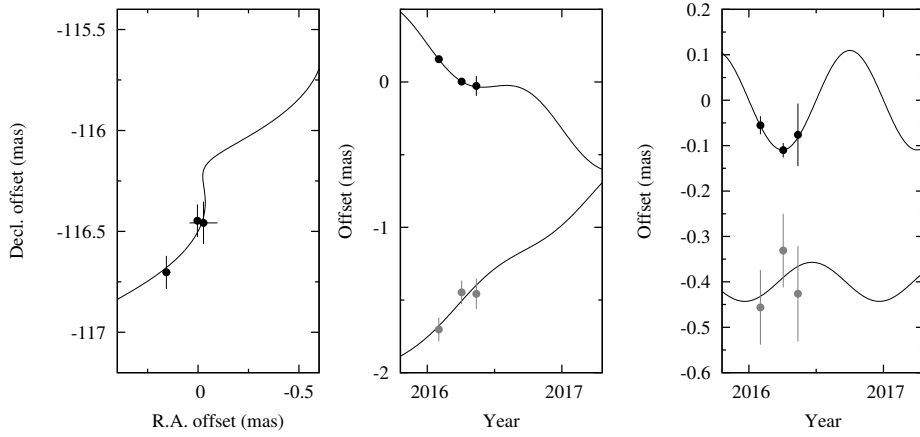




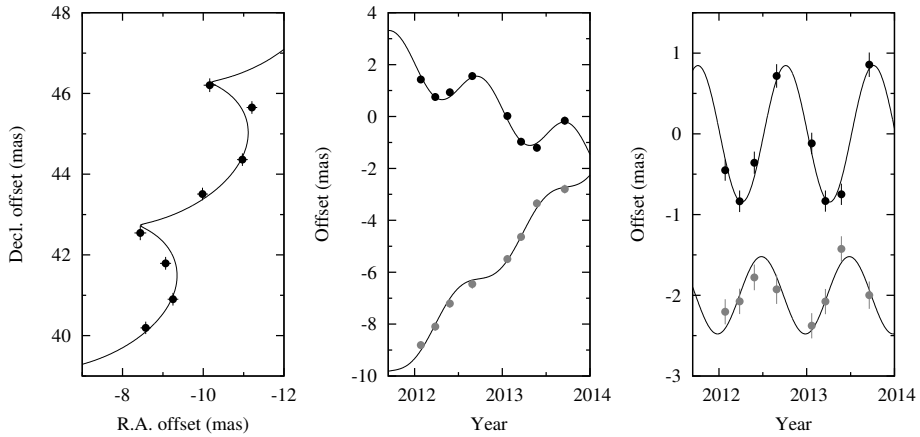
IRAS06501+0143  
 06h52m45.3205s +01d40'23.048"  
 $\pi = 0.283 \pm 0.108$  mas ( $D = 3.53^{+2.18}_{-0.98}$  kpc)  
 $(\mu_x, \mu_y) = (-0.72 \pm 0.15, 0.66 \pm 0.19)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.113, 0.150)$  mas



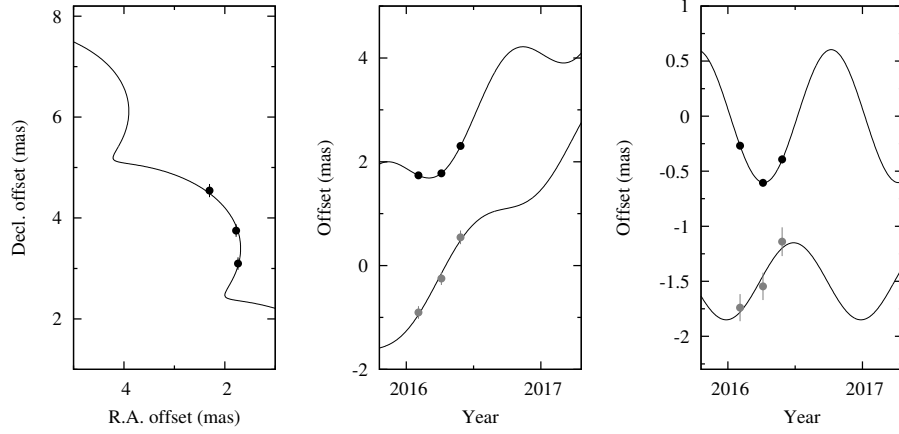
WB886  
 06h47m13.3344s +00d26'05.920"  
 $\pi = 0.110 \pm 0.130$  mas ( $D = 9.1^{+X.X}_{-X.X}$  kpc)  
 $(\mu_x, \mu_y) = (-0.58 \pm 0.35, 0.77 \pm 0.53)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.010, 0.080)$  mas



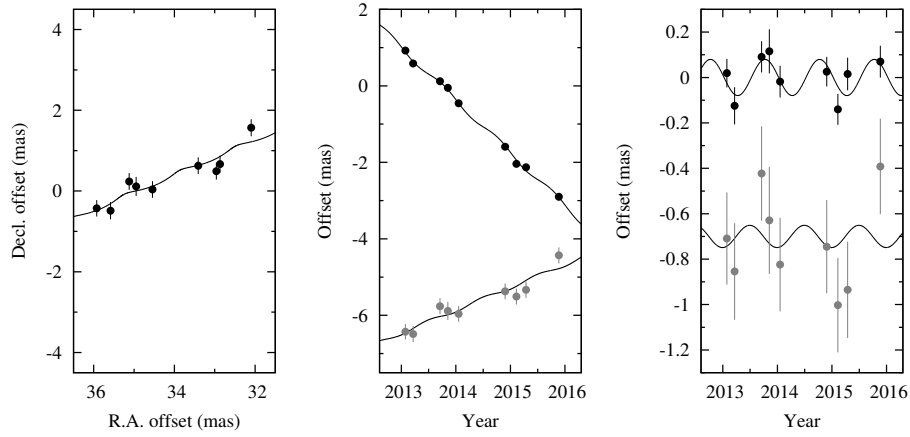
IRAS07024-1102  
 07h04m44.3900s -11d07'12.450"  
 $\pi = 0.850 \pm 0.070$  mas ( $D = 1.18^{+0.11}_{-0.09}$  kpc)  
 $(\mu_x, \mu_y) = (-1.76 \pm 0.09, 3.54 \pm 0.10)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.131, 0.155)$  mas



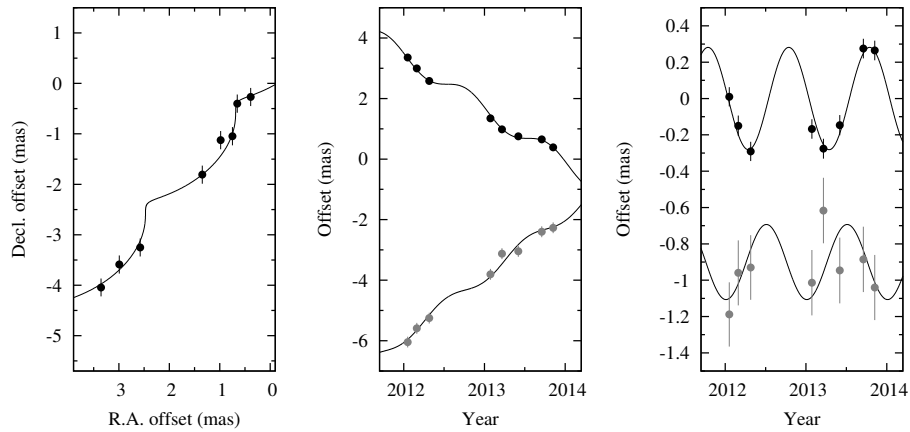
G226.27  
 07h14m26.9333s -12d03'58.884"  
 $\pi = 0.609 \pm 0.070$  mas ( $D = 1.64^{+0.21}_{-0.17}$  kpc)  
 $(\mu_x, \mu_y) = (2.22 \pm 0.15, 2.72 \pm 0.59)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.005, 0.119)$  mas



IRAS07207-1435  
 07h23m01.8407s -14d41'32.807"  
 $\pi = 0.080 \pm 0.034$  mas ( $D = 12.5^{+9.2}_{-3.7}$  kpc)  
 $(\mu_x, \mu_y) = (-1.38 \pm 0.03, 0.60 \pm 0.08)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.062, 0.228)$  mas

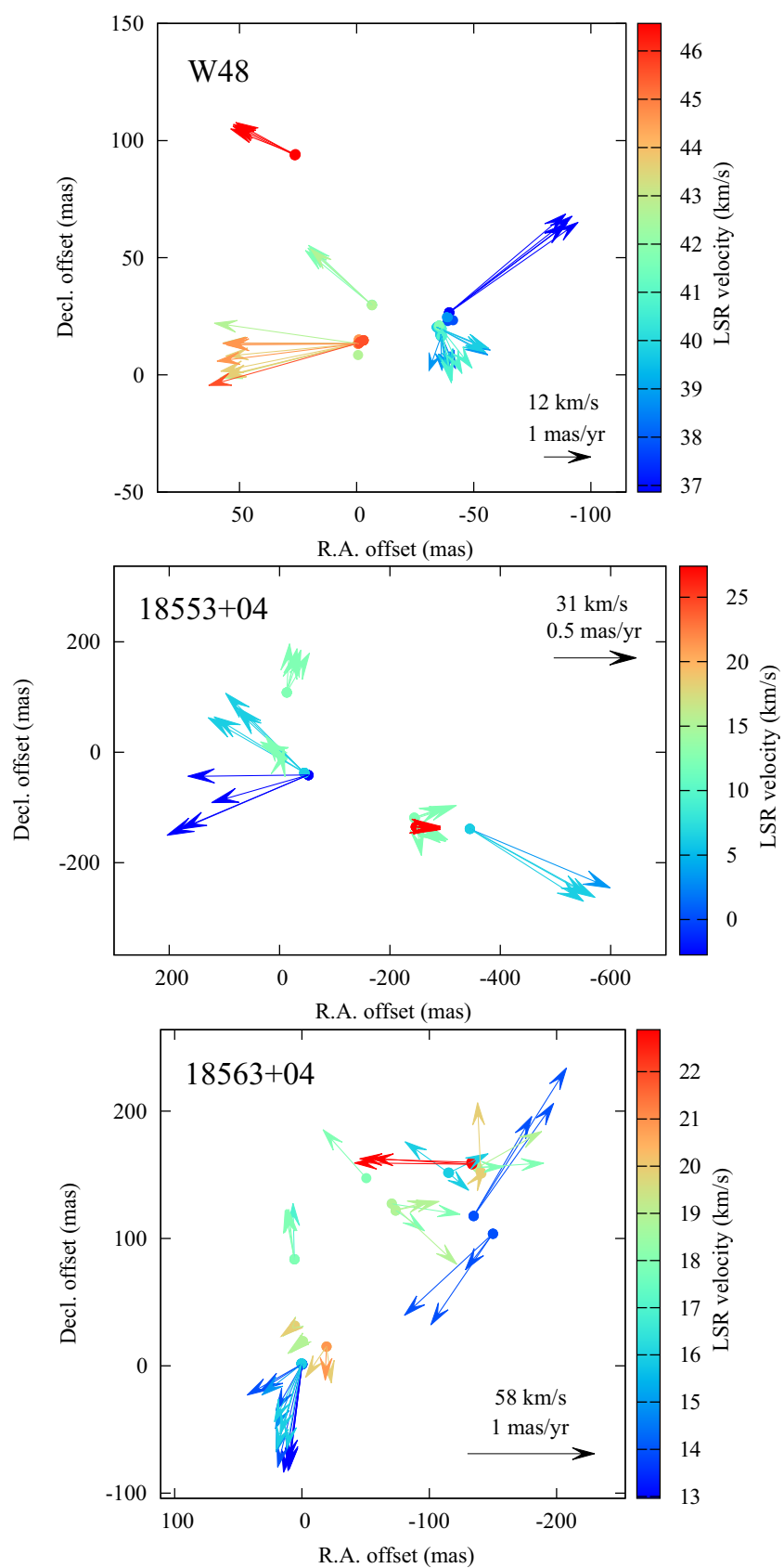


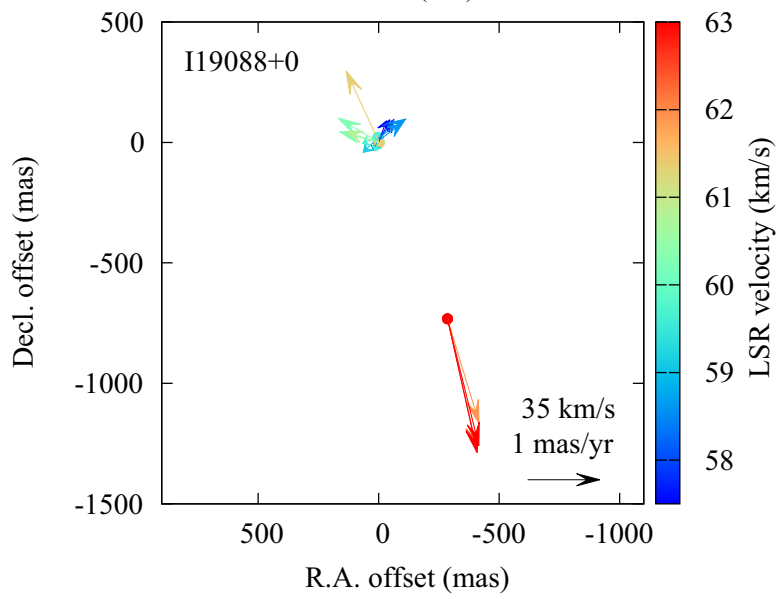
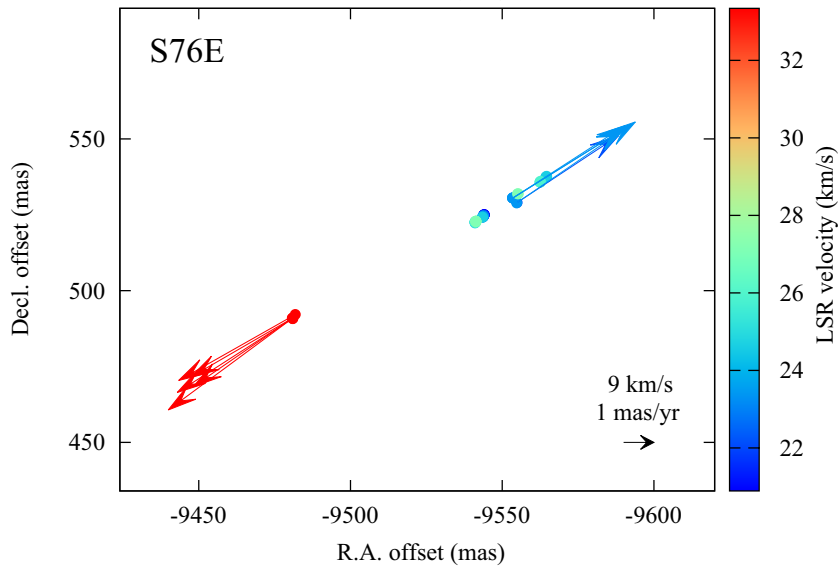
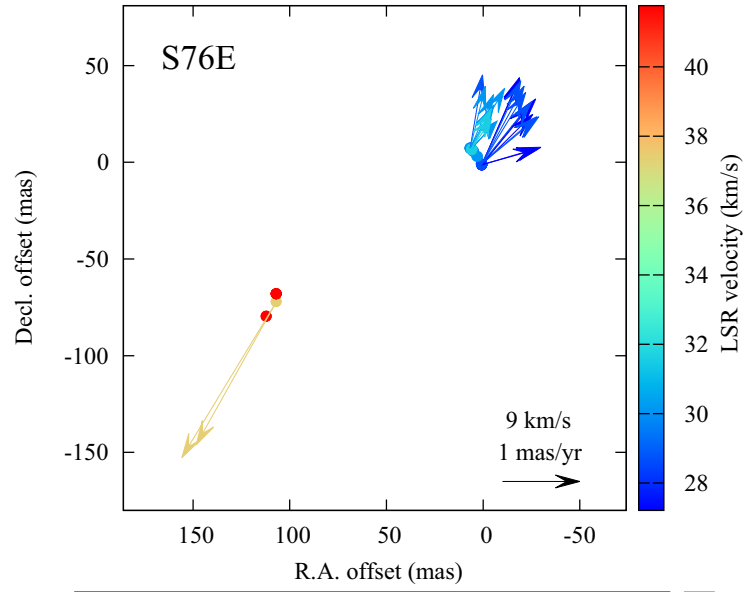
IRAS07427-2400  
 07h44m51.9368s -24d07'42.811"  
 $\pi = 0.286 \pm 0.035$  mas ( $D = 3.50^{+0.49}_{-0.38}$  kpc)  
 $(\mu_x, \mu_y) = (-1.79 \pm 0.04, 2.01 \pm 0.10)$  mas/yr  
 $(\sigma_x, \sigma_y) = (0.054, 0.180)$  mas

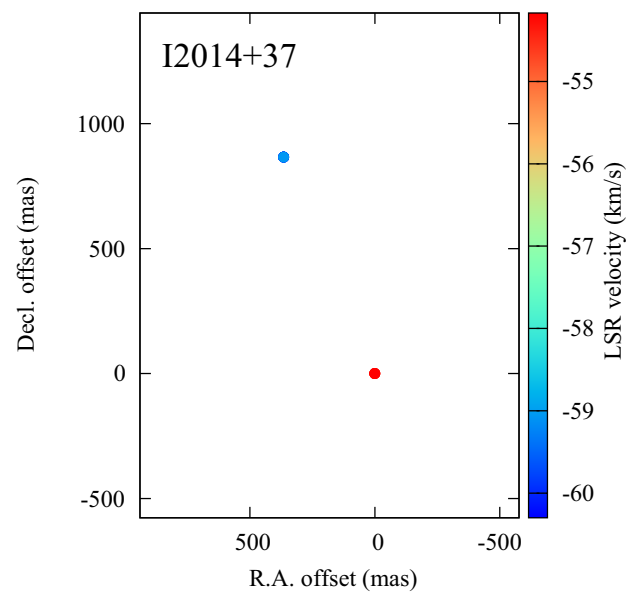
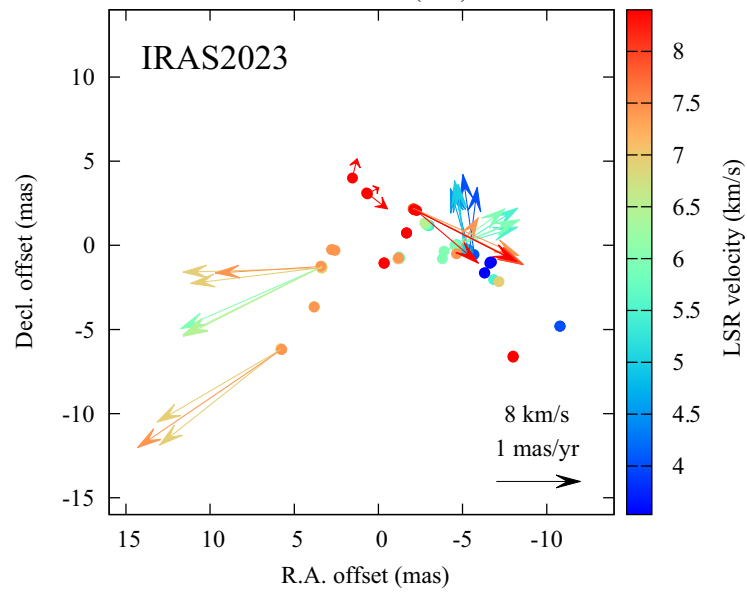
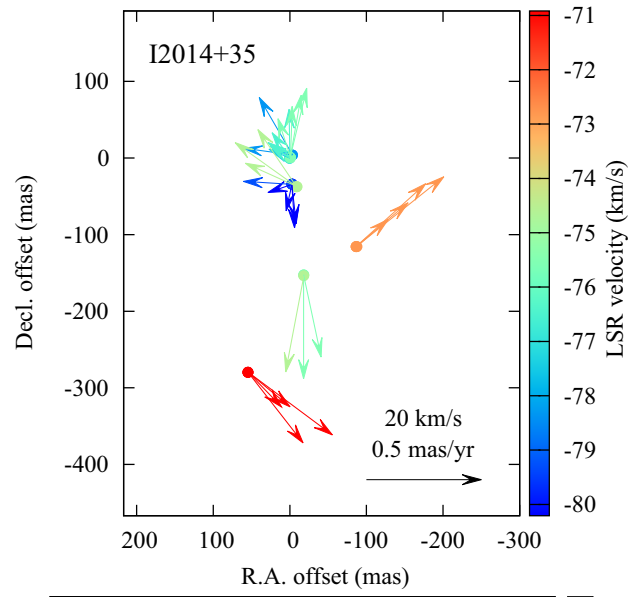


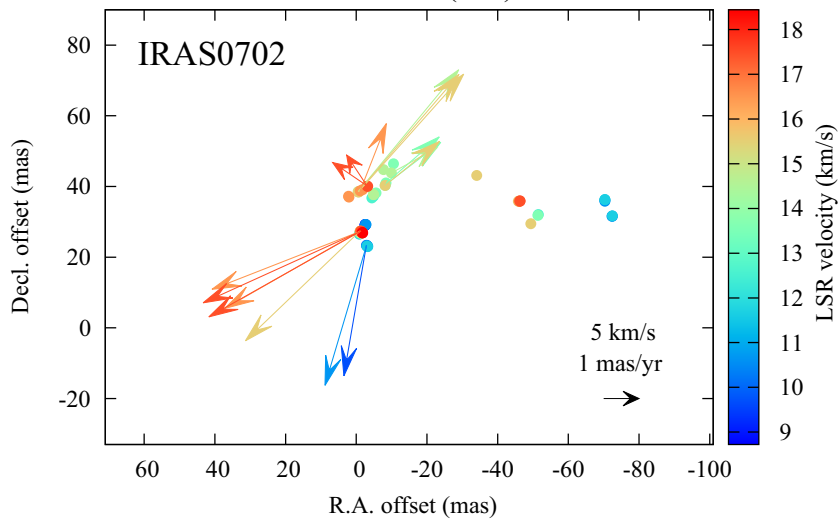
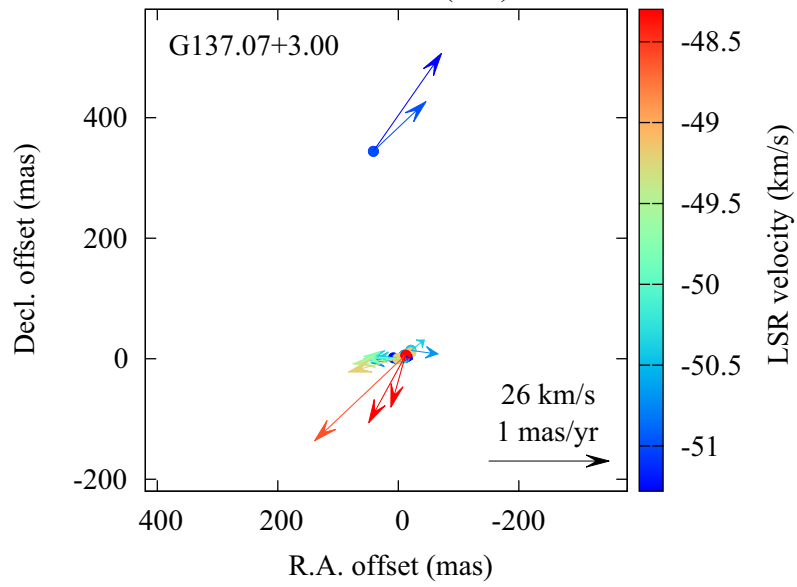
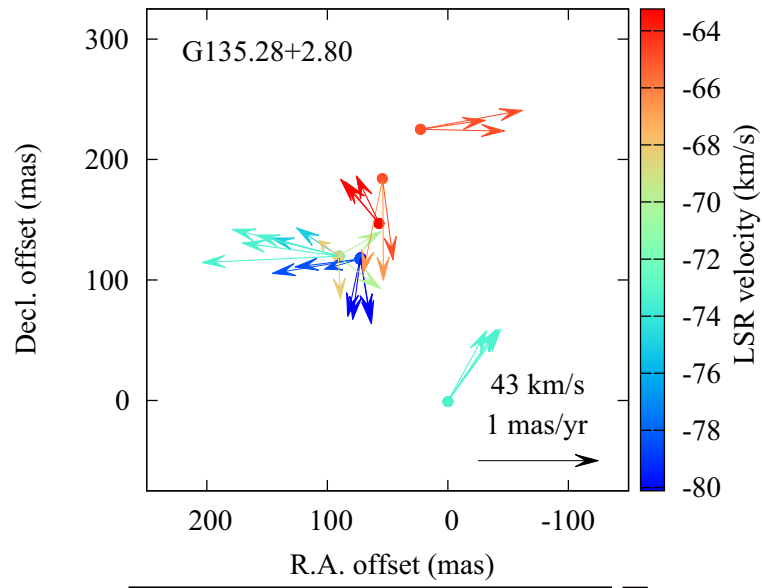
### 3 内部運動

12 天体のメーザーの内部運動を示す。









## 4 銀河系内での分布と銀河回転曲線

視差測定済みの天体数は、昨年の約 50 天体から 86 天体 (SFR が 70 天体、AGB が 16 天体) となった。Bessel(VLBA/EVN) を含めると 159 天体となった。視差測定済みの SFR143 天体の銀河系内での分布と銀河回転曲線を示す。

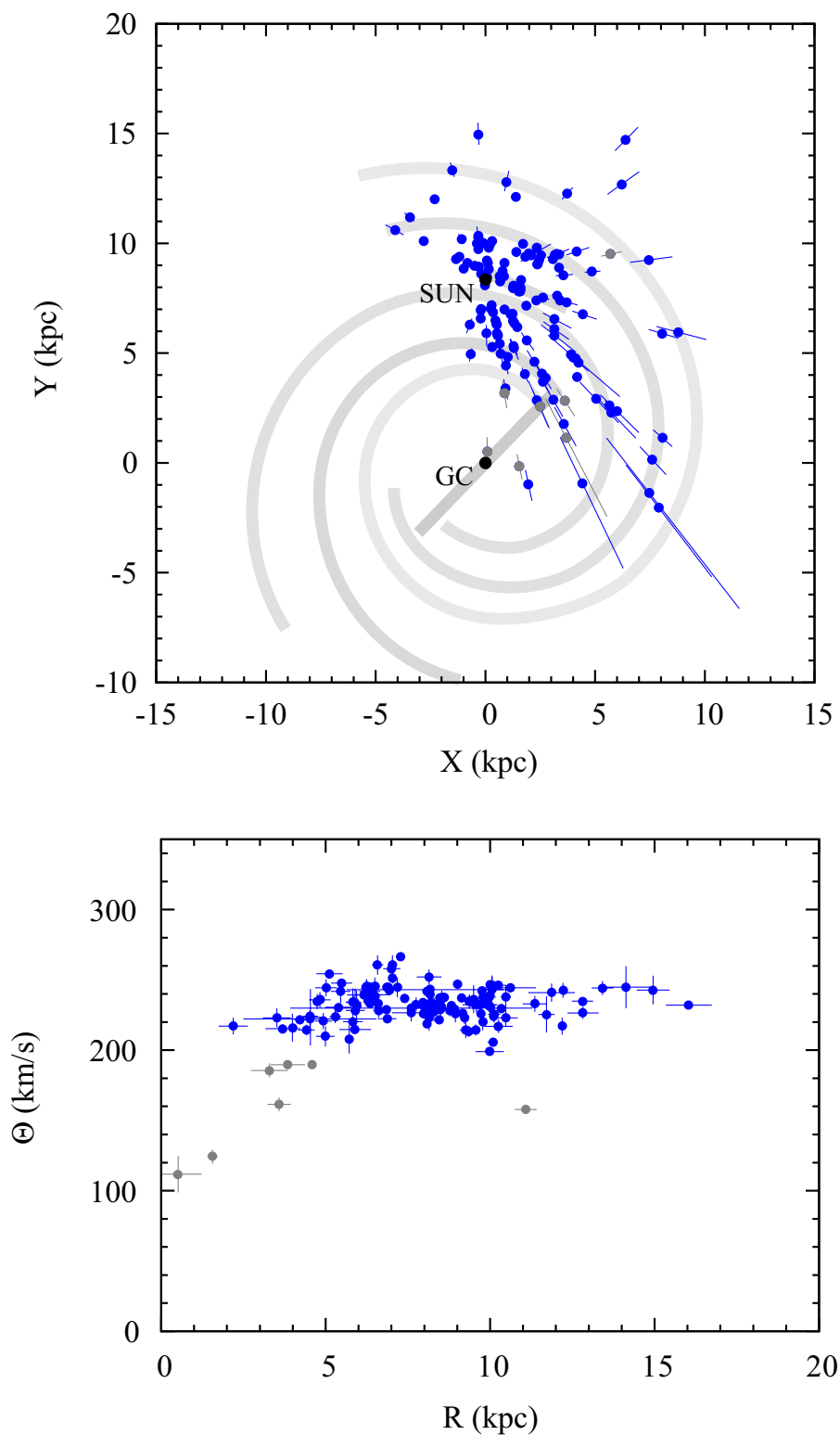


図 1: 銀河系の分布 (上) と銀河回転曲線 (下)。

## 5 銀河系パラメータの推定

視差測定済みの SFR143 天体のうち、bar 付近の非円運動の大きい 7 天体を除いた 136 天体を用いて、銀河定数 ( $R_0, \Omega_0$ )、rotatin curve index ( $\alpha$ )、天体の非円運動 ( $U_s, V_s, W_s$ ) を求めた。結果を表 2 に示す。銀河定数の精度は Honma et al. (2012) の 52 天体時の約 5% から約 3% へ向上した。

表 2: 銀河系パラメータ

$R_0$	$8.26 \pm 0.24$ kpc
$\Omega_0$	$28.84 \pm 0.37$ km s <sup>-1</sup> kpc <sup>-1</sup> ( $\Theta_0 = 238 \pm 8$ km s <sup>-1</sup> )
$\alpha$	$0.011 \pm 0.015$
$U_s$	$4.00 \pm 1.09$ km s <sup>-1</sup>
$V_s$	$-5.33 \pm 1.01$ km s <sup>-1</sup>
$W_s$	$-0.61 \pm 0.85$ km s <sup>-1</sup>

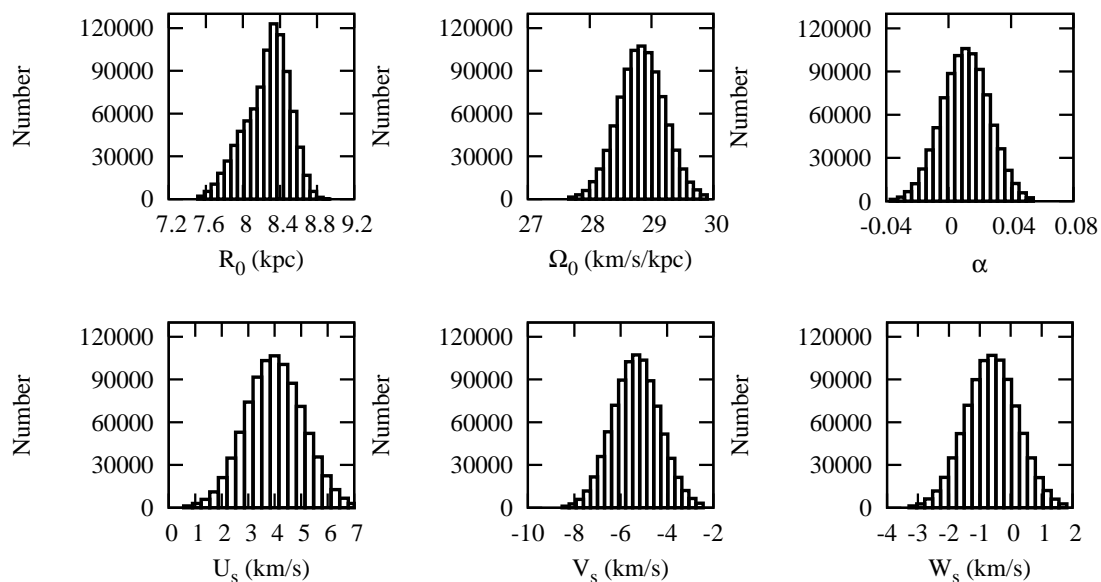


図 2: ヒストグラム